

# Diameters in the Beta Pic Moving Group Measured with CHARA/CLASSIC

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Observed in 2010 and 2011 in observing time assigned through the NOAO

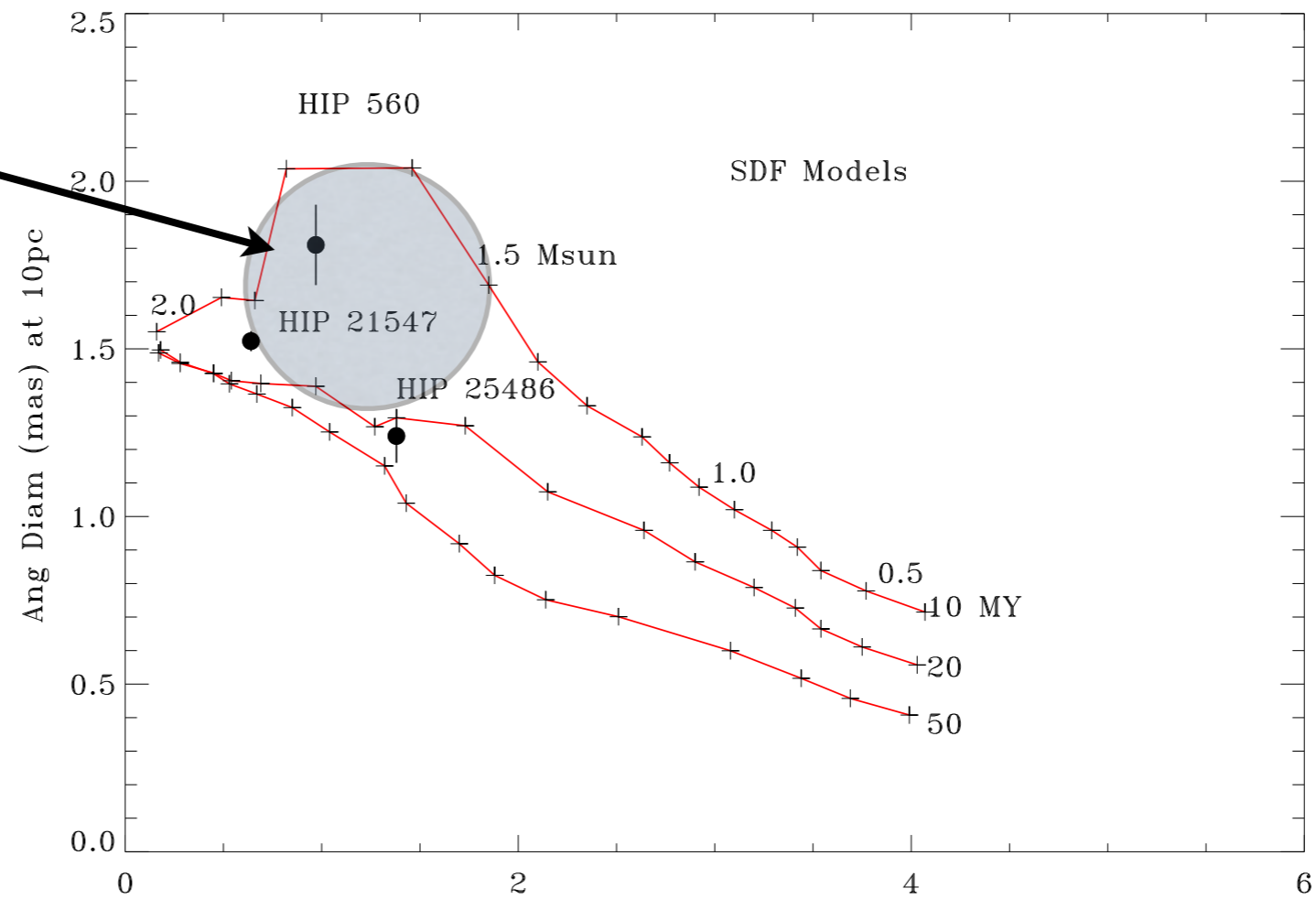
2010 results published in ApJ 743, 158 (2011)

Goal: To determine the ages of young stars by measuring their diameters as they contract to the main sequence.

“Sweet Spot” for observations, F and G spectral type stars

CAUTION:

Young A, F stars rotate fast!

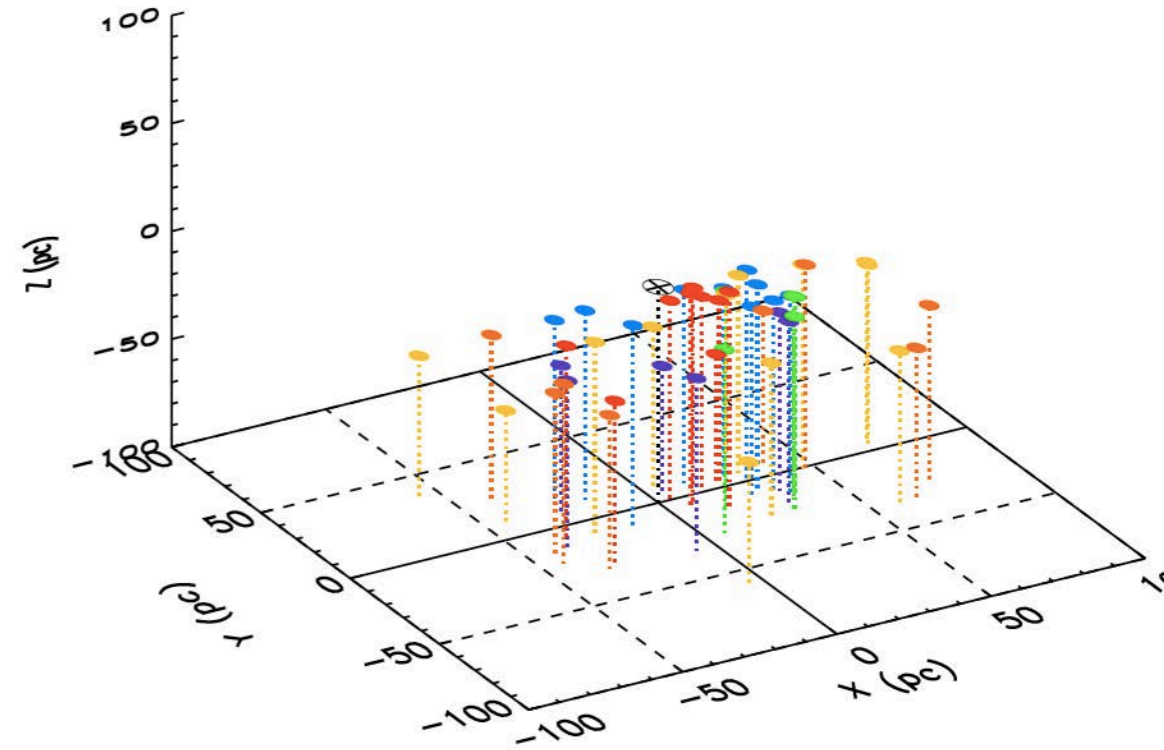
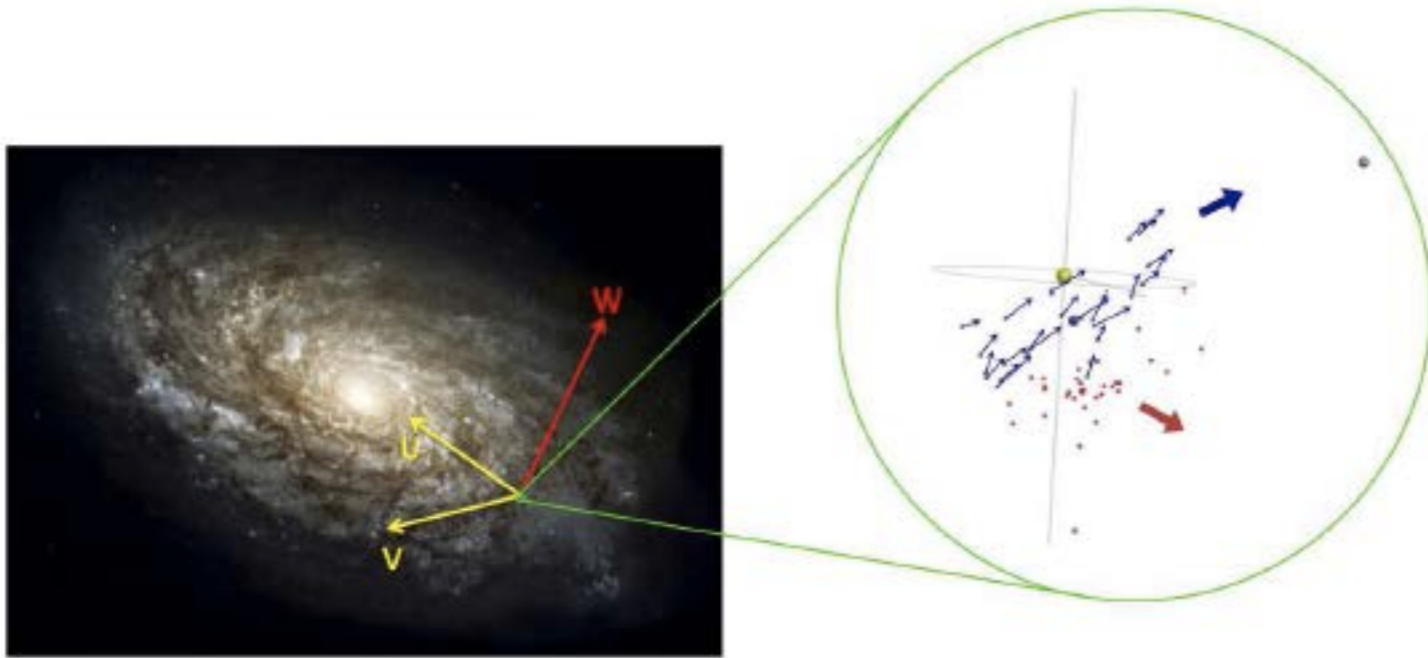


Models calculated by different groups agree well

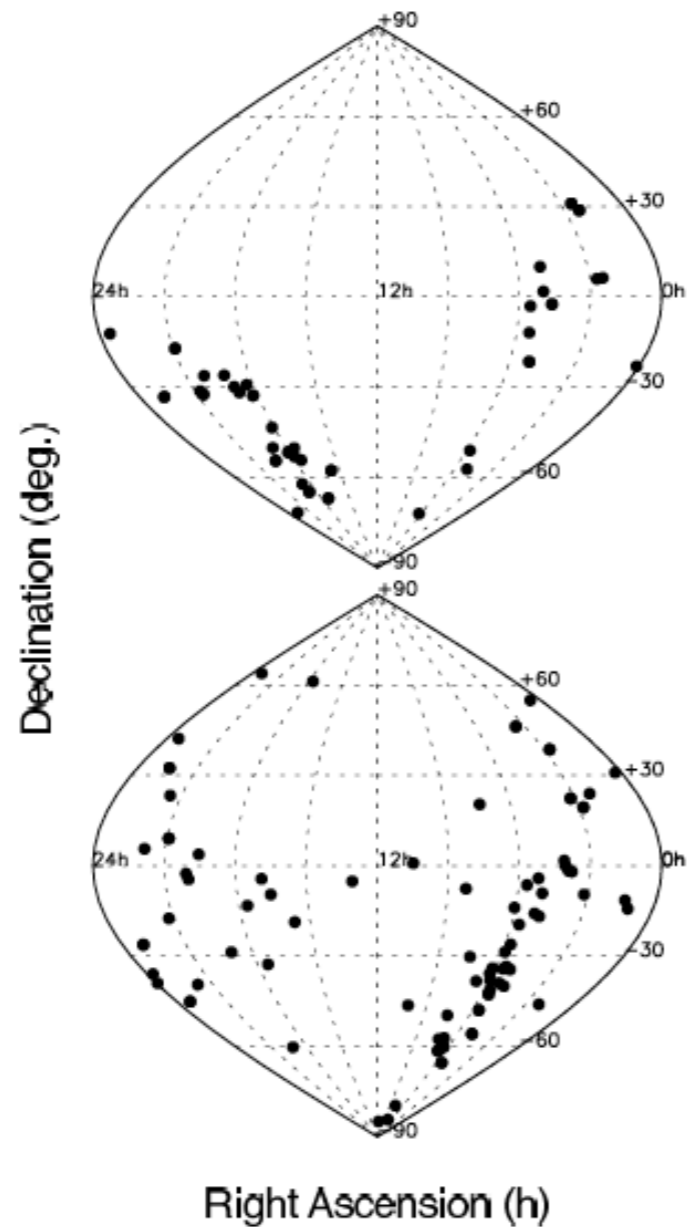
Models differ at  $M < \sim 1$  Msun

## Why the Beta Pic Moving Group?

1) Its stars are young (10-20 My) and nearby (median  $d=35$  pc)



2) Some of its stars are observable in the north



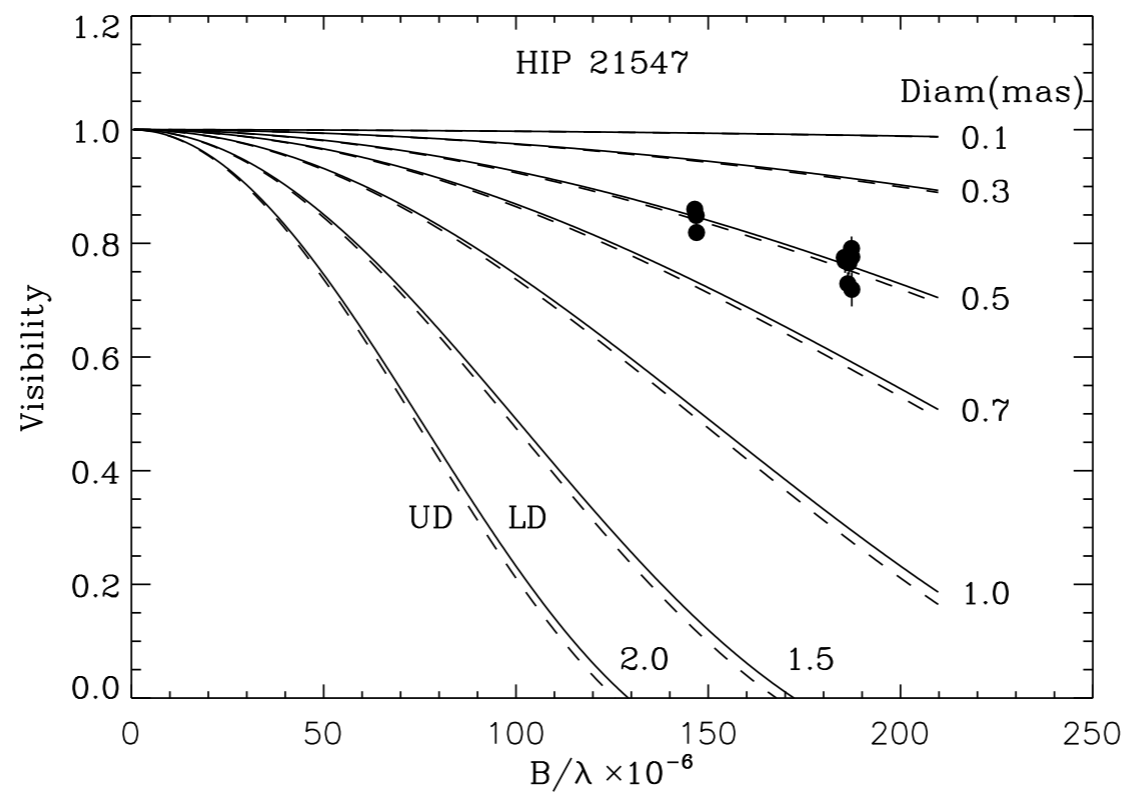
Beta Pic Moving Group

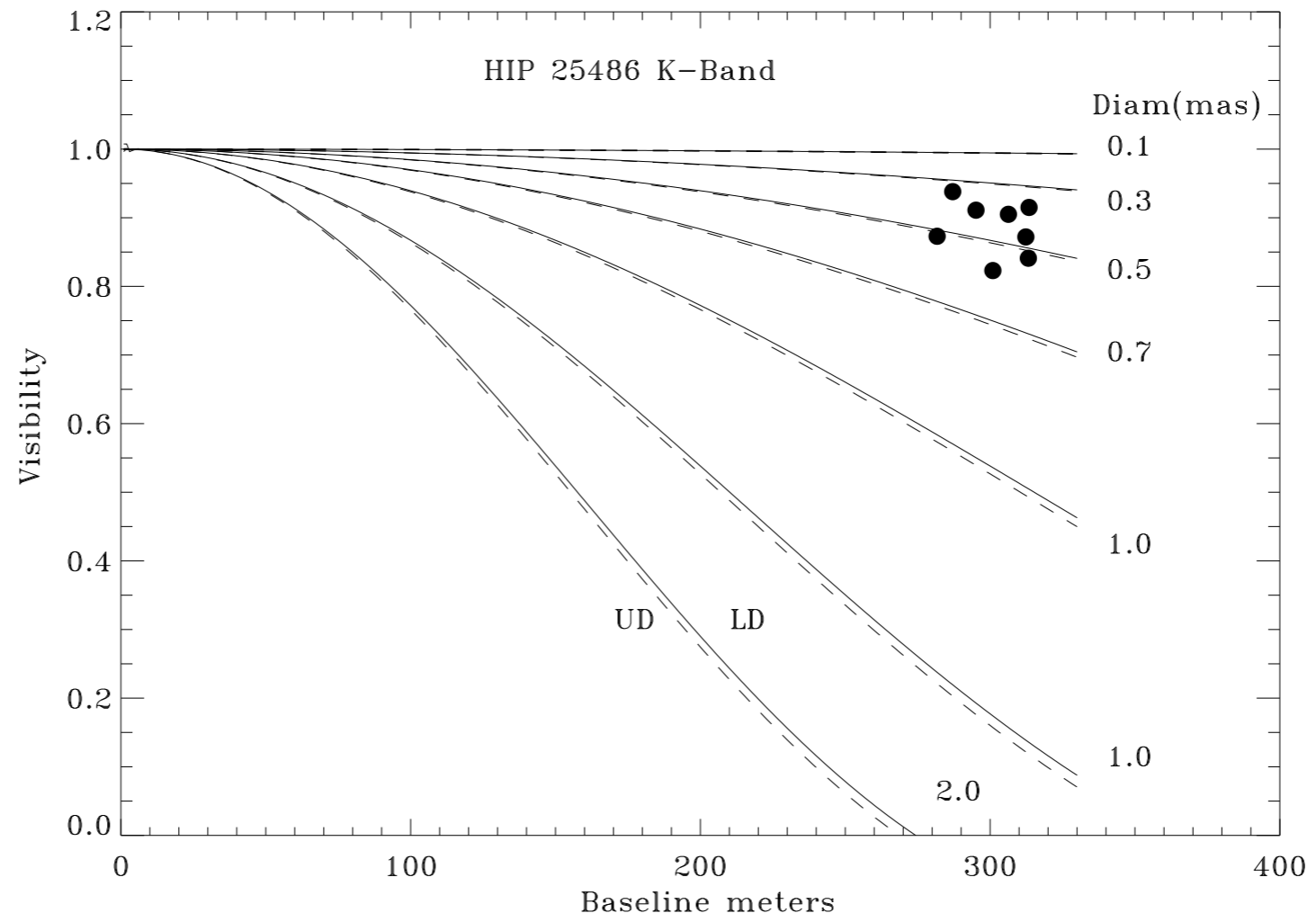
AB Dor Moving Group

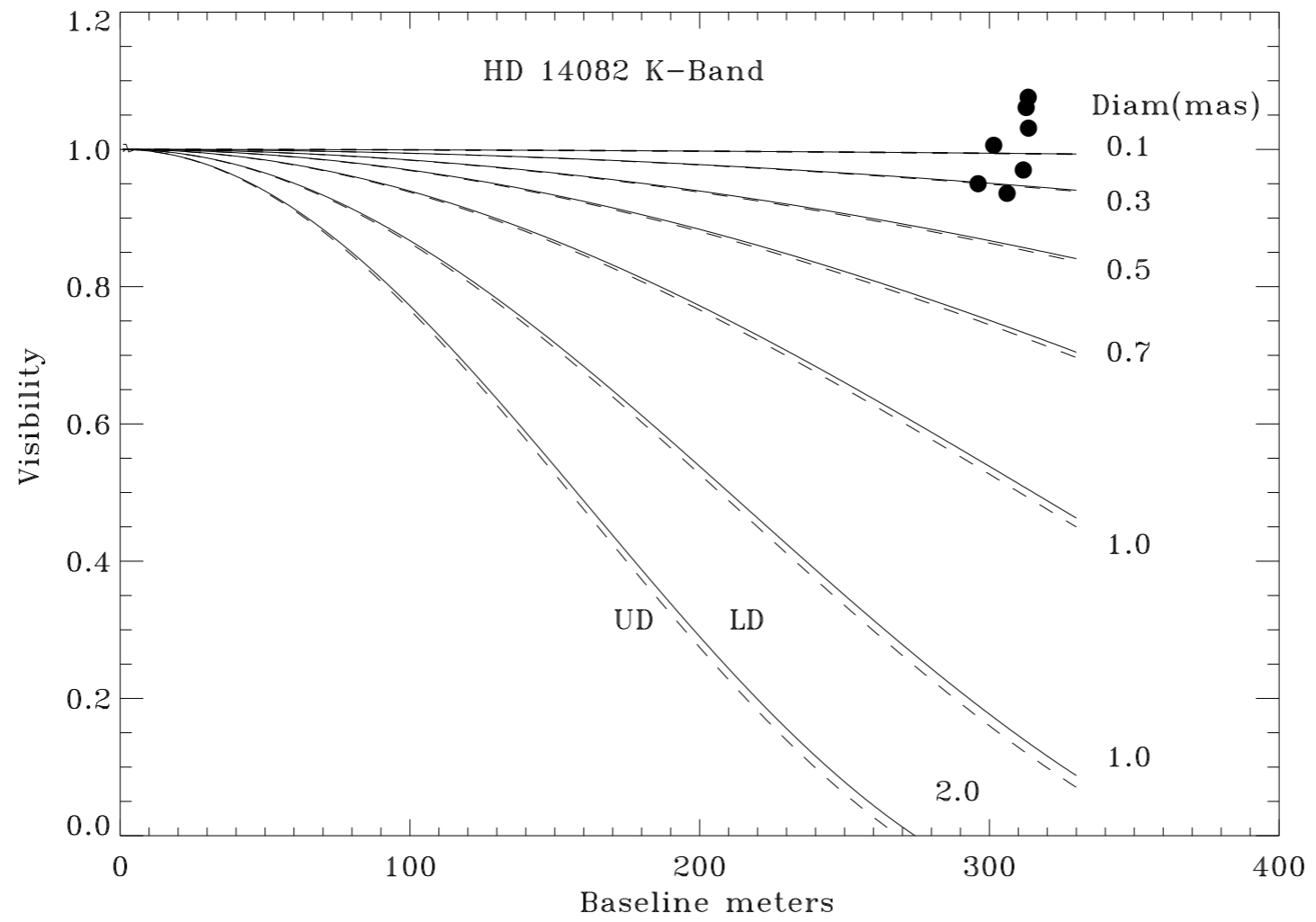
## Stars Observed:

Star	SpTy	K	Dist pc	Diam Msrd mas	Predicted <sup>*</sup> McCarthy/White
HIP 560	F3V	5.2	39.4±0.4	0.46±0.03	0.37
HIP 21547	F0V	4.5	29.4±0.3	0.52±0.01	0.52
HIP 25486	F8V	4.9	27.0±0.4	0.46±0.03	[0.40]
HIP 10680 (HD 14082A)	F5V	5.8	34.5±3.5	<0.36 3-sigma	0.30

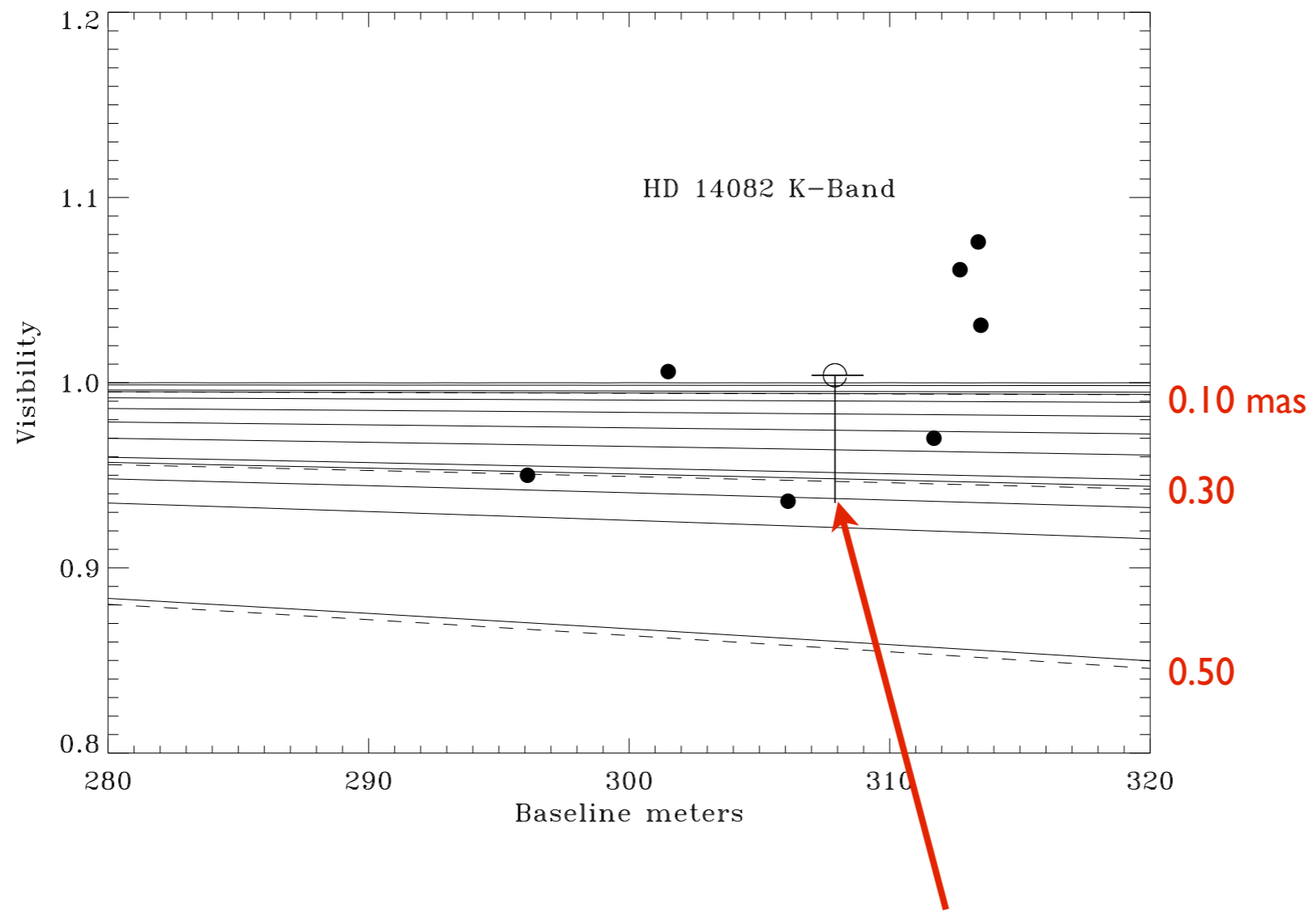
From empirical  $L, T_{\text{eff}} \Rightarrow R_{\text{star}} + \text{Hipparcos parallax}$ , AJ, in press

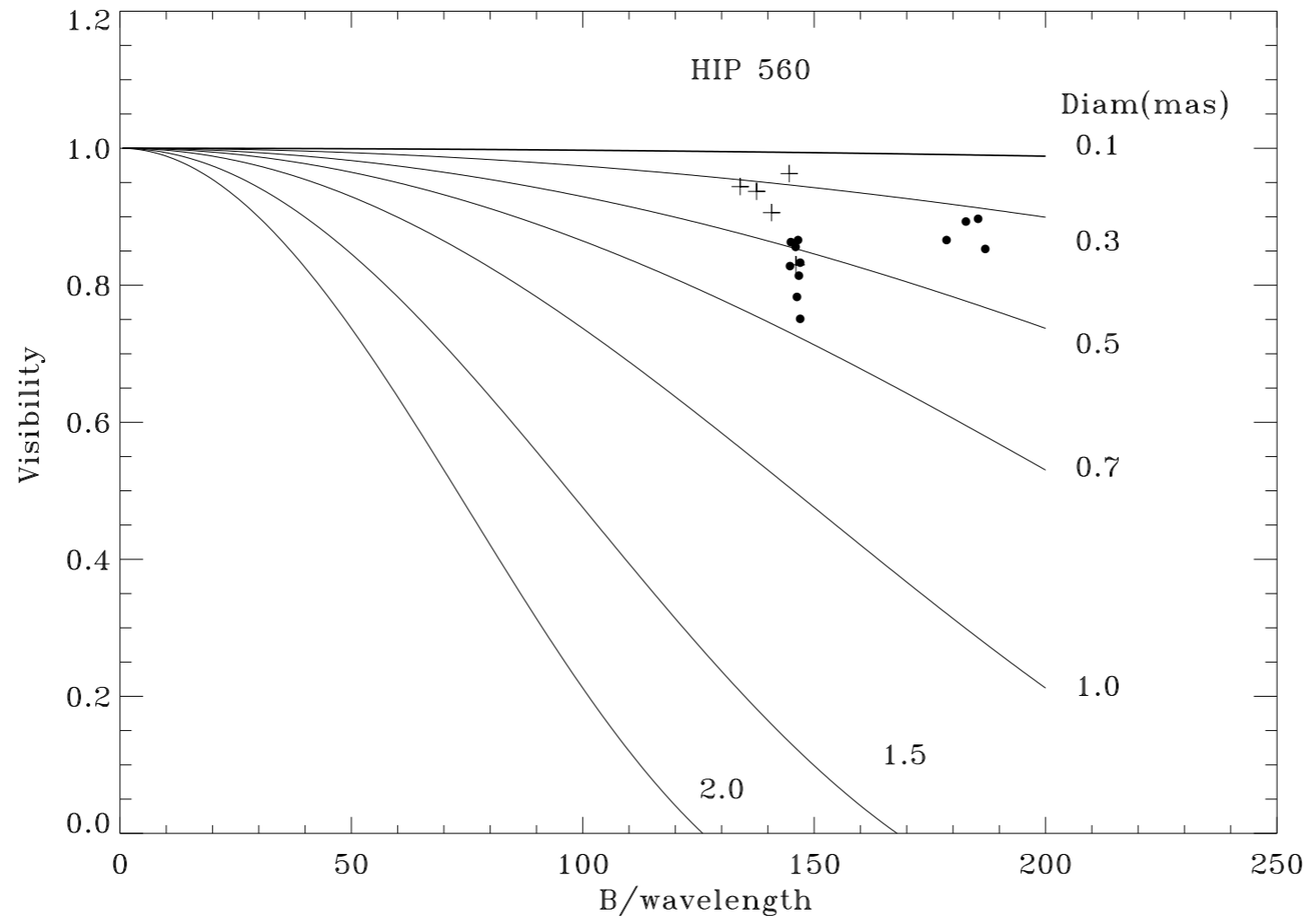






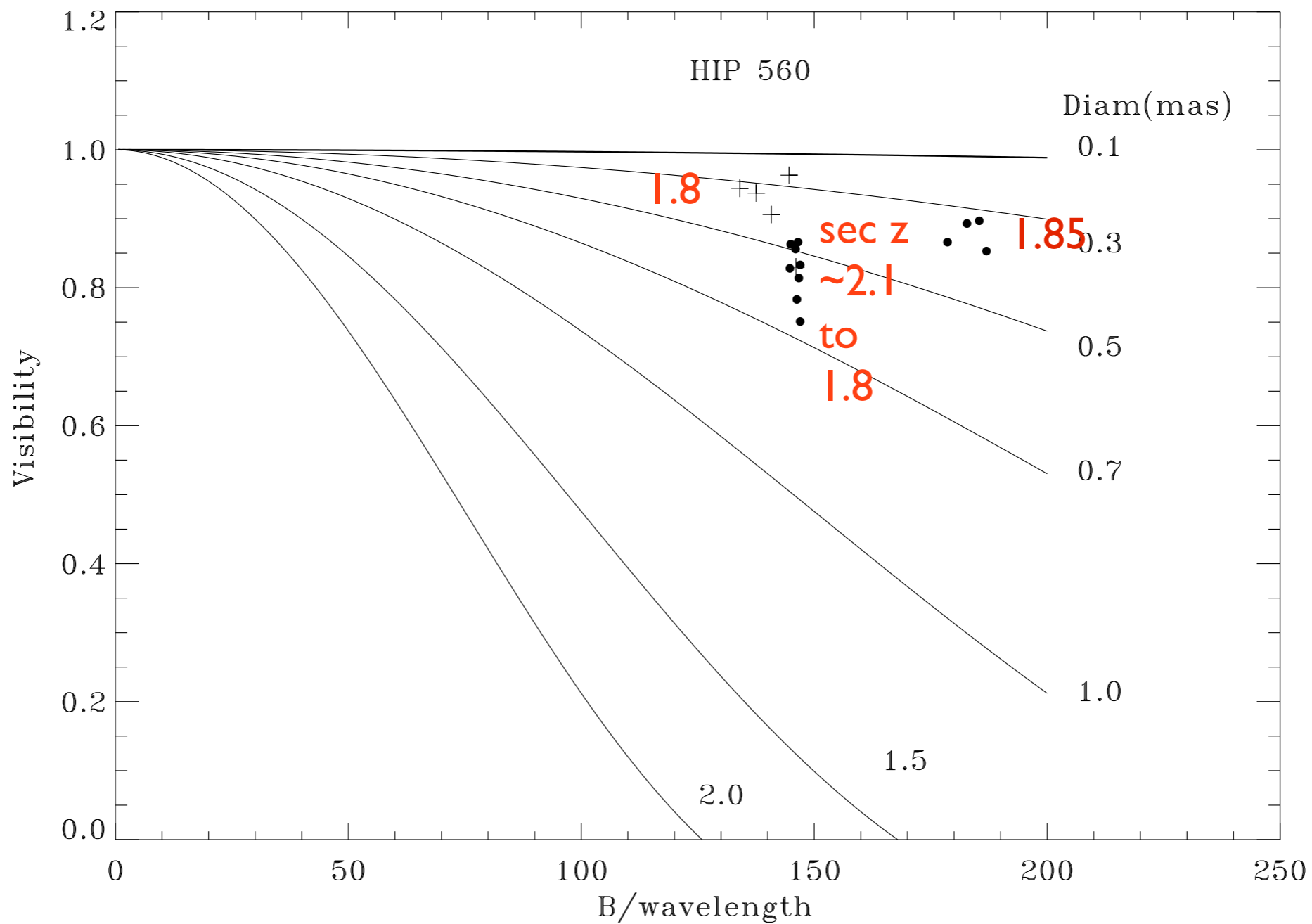




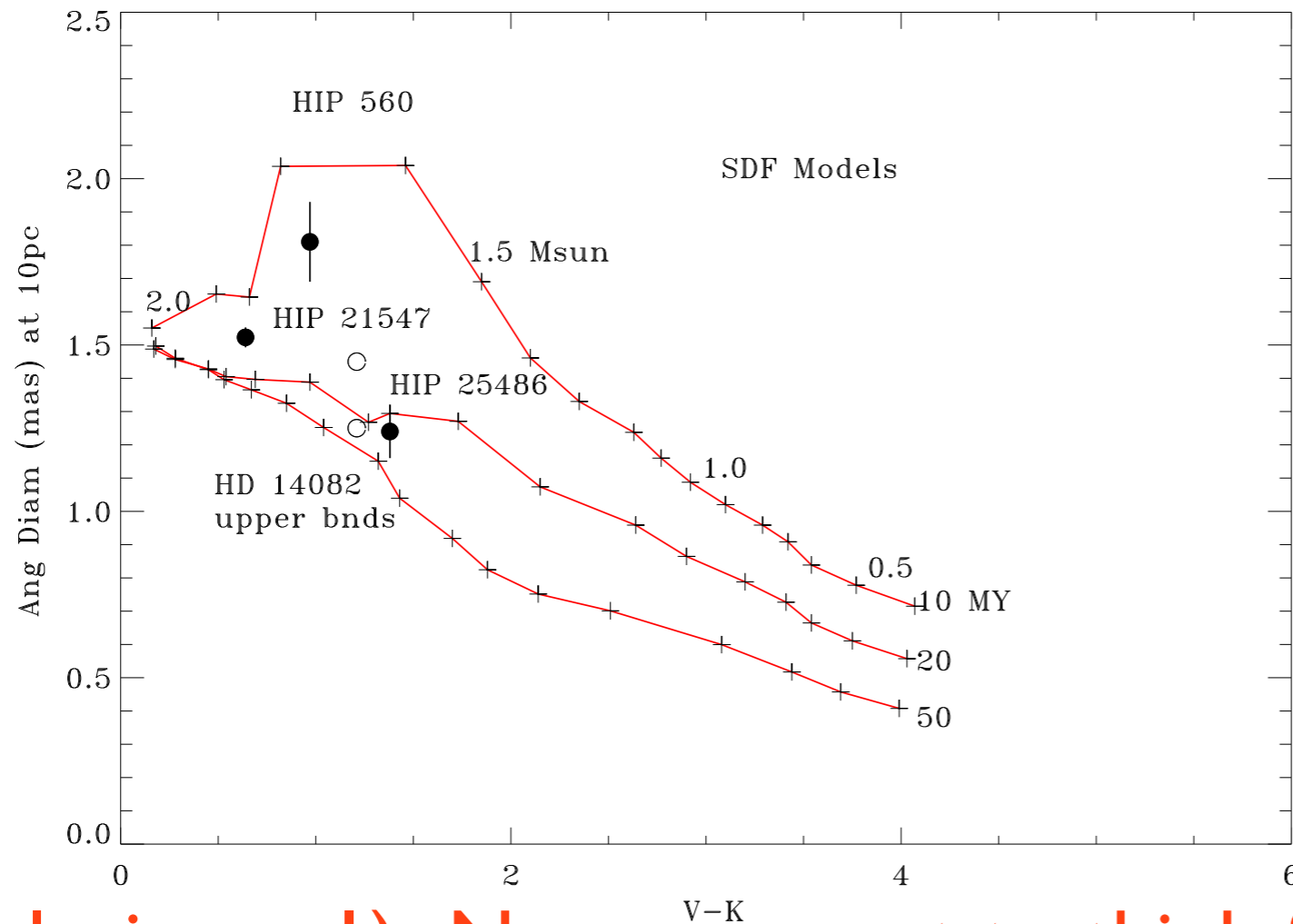


• K and H, 9/2010

+ K, 11/2011



## Results:



Star

Age My

HIP 560

13 $\pm$ 2

HIP 21547

15 $\pm$ 2

HIP 25486

25 $\pm$ 10-6

HD 14082

>16

- Conclusions:
- 1) No reason yet to think formation not isochronous
  - 2) BPMG Age probably 15-20 My
  - 3) Need to find K, M stars at  $D < 10$  pc in the north

Lessons: 1) Beware large airmass

