



# Perspectives on a new visible 6T instrument on CHARA

Nice, 16 Mar 2016



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NASA Exoplanet Science Institute



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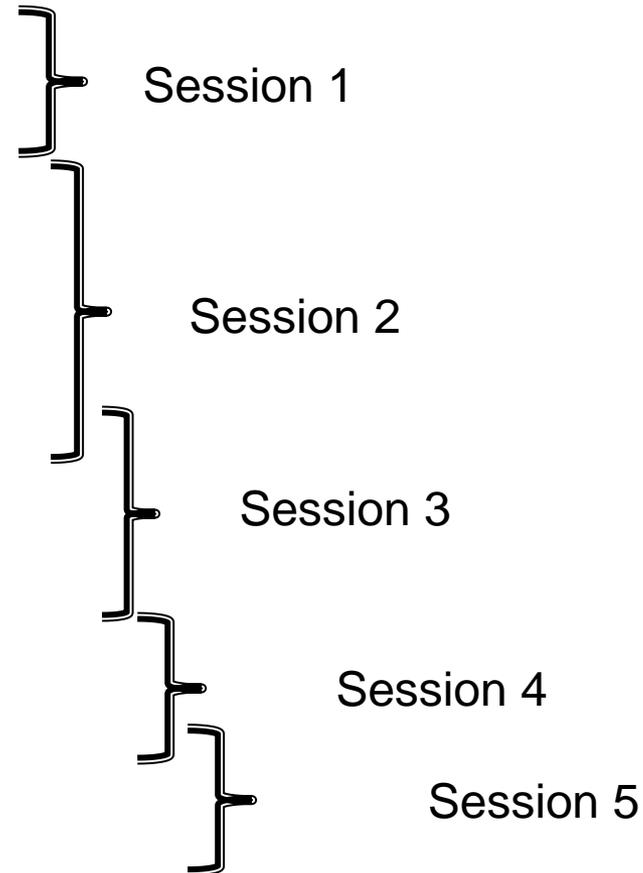
# Introductory remarks

- 8 years of VEGA and PAVO: limitations identified
- AO perspectives and progresses on the FRIEND demonstrator.
- Initial idea presented during CHARA Meeting in Atlanta last year
- Kick-off meeting in Nice on 21 Sep 2015 + follow-up meeting 11 Jan 2016 → a plan in ~20 actions for the conceptual study.  
→ this meeting



# Main actions

- Science program
- Detector issue
- Beam combination
- Fiber injection
- Fringe tracker
- Software
- General organization





# SESSION 1 – SCIENCE WITH VISIBLE BEAM COMBINERS



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# A short summary of the previous steps

White book for a visible interferometric instrument: Ph. Stee et al.

- Initial start at EII level by Olivier in 2012/2013
- Working groups by Philippe in 2014
- Various ESO meetings end of 2014/early 2015
- Meeting and publications in 2015

VLTI perspectives and CHARA framework



# OPPORTUNITY FOR LARGE PROGRAMS





# Atlanta 2015



CHARA 2015 Annual Science Meeting - ATLANTA

## We need a visible 6T beam combiner

- Excellent opportunity for **fundamental parameter** of stars survey or **'legacy' program**. Already started yes but could be extended thanks to visible wavelengths, sensitivity and 'fast' observations with 6T.
- Also excellent case for **imaging and spectral imaging**
- Convergence of various efforts
  - PAVO, VEGA: visible and spectro-interferometry
  - MIRC: imaging and 6T operations
  - FLUOR: high precision
  - AO@CHARA
- Science preparation and technology developments



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# White book cases for LP

- Surface-brightness-color relations
- Masses
- Fundamental parameters as a function of SpTy
  - Radius
  - Effective Temperature
  - Limb darkening

➔ FROM EMPIRICAL MODELS TO MEASURES

- Exoplanets host stars
- Asteroseismic targets
  
- Original and unique support to a large panel of space missions
- *Doug's opening remark:*
  - *single stars: Gaia, TESS, CHEOPS, PLATO;*
  - *binaries: stellar masses and stellar evolutionary outcomes*



$$5 \text{ years} * 50 \text{ nights} * 8 \star = 2000 \text{obj}$$

- 8 stars / nights
  - $8 \times (\text{Cal1} + \text{Obj} + \text{Cal2}) * 2 = 48 \text{ o/n}$
- 2015: 187n, 4699o → 16o/n. Factor 3 missing
- But
  - AO → 75% of ‘good seeing’ conditions
  - 6T means  $15V^2$
  - Low spectral resolution and number of channels  $\lambda$ -coverage
- Probably feasible:
  - PAVO/FLUOR experience of fast shift
  - Queue scheduling and optimization



# SESSION 2: INSTRUMENTATION



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# Main open questions

- Detector: OCAM2 – Nüvu
- Beam combiner
  - FRIEND characterization (MAM)
  - Various combining schemes
  - IO tests
- Fibers



# SESSION 3: INTERFACES WITH AO AND FT





# Main open questions

- Quality of injection behind a partial AO?
  - Additional T/T?
  - Interfaces with the various CHARA loops?
- Need for a 6T coherence/fringe tracker
  - Science limits
  - Coherencing+ Coherent processing?



# SESSION 4: CONTROL, OPERATION, DRS





# Main open questions

- Principles for the instrument control software  
→ CHARA standard
- Estimation of data flow, perspectives of archiving
- Data processing principles?



# SESSION 5: DISCUSSION





# Introduction

- Nice+Grenoble+Lyon initiative
- Based on VEGA+FRIEND
- But should not be a VEGA-like instrument
  - ... but a full CHARA one
  - ... and if possible and interesting: 1 for CHARA and 1 for VLTI
- Timeline for CHARA >2018?
- Timeline for VLTI >2020?
  
- Funding request to be prepared for next summer
- Partnership to be defined
  - many CHARA interfaces
  - Experience from PAVO, example of VISION
  - Synergy with MYSTIC/MIRCx for Group delay
  - FT-Gravity

# Today's main points of conclusion?

- Science
  - Imaging/spectral-imaging
  - High-efficient  $\theta$  machine
    - Sensitivity (detector, AO, FT)
    - Simple instrument with integrated pipeline, very few modes (2-3 maximum)
- Concept
  - Technology ready
  - Bulk/IO: choice to be sensitivity-driven
  - Multi-axial, dispersed fringes
- Interfaces (important for sensitivity and reliability)
  - Pupil and image trackers, LDC
  - AO and fiber injection
  - NIR for group-delay/fringe tracker: residuals versus magnitude/ $r_0$ /dispersion...
- Control/operation/DRS
  - To be built on previous experience (see JMC talk) and CHARA integrated
  - More automatic processes towards final products