



SPICA *S*tellar *P*arameters and *I*mages with a *C*ophased *A*rray

A new visible 6T spectro-combiner for CHARA

<https://lagrange.oca.eu/fr/spica-project-overview>

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& S. Kraus, S. Lacour, J.B. Le Bouquin, J. Monnier, & S. Albrecht
& the CHARA staff (special thanks to Judit!)



Origins of the project

- **Main scientific idea**
 - Growing interest for fundamental parameters of star: Radius ($\theta + \pi$), Effective temperature ($\theta + F_{\text{bol}}$) for exoplanet host stars, distance scale (with SBCR), asteroseismology. Analysis of the stellar noise due to “activity”
 - Importance to develop an homogeneous and large survey to cover a wide region of the HR diagram.
- **Opportunity**
 - Development of the AO program on CHARA
 - Identification of the main limitations of VEGA (detector, multimode regime) and solutions (FRIEND)
 - Huge progress on EMCCD, single mode fibres, integrated optics device (Gravity Fringe Tracker)
- **Top level requirements**
 - Quality of measurements: Spatial filtering, 6T operation
 - Sensitivity up to mag 8 (10 ideally): low spectral resolution, fringe tracking, optimised injection
 - Survey mode with optimised night strategies
 - Combination of R, (J) H K data



SPICA in a nutshell

- A H-band 6T-ABCD fringe sensor aiming at performing a group delay and phase delay tracking of the fringes.
- A All-In-One 6T combiner (600-900nm) with three dispersion

MODES	Nb of SpCh	SpCh	Spectral Band	MagLim $V^2=0.6$	MagLim $V^2=0.6 + FT$	MagLim Vdiff	MagLim Vdiff+FT
LR: R=140	60	~5nm	300nm	6.5	11.5		
MR: R=4400	500	0.17nm	85nm			5.5	9.5
HR: R=13000	500	0.06nm	29nm			4.5	8.5

MagLim: V^2 : SNR=10, 10mn of integration, for one spectral channel

Vdiff: [SNR=10, $\sigma\phi < 5^\circ$], $V^2=0.6$ in the reference channel, 30mn of integration, for one spectral channel

SNR calculator based on FRIEND calibration (Martinod+2018), CHARA-AO hypothesis **SR=5%**, SPICA estimations
Validation to be done on sky beginning of 2022

Scientific Requirements: θ_{LD} and $(\theta_{LD} + LD)$ measurements

θ_{LD} only with hypothesis on the LD law

$$\sigma\theta/\theta=1\%$$

θ_{LD} and LD from SPICA

Dwarfs	Challouf			Salsi-1			Salsi-2		
SpTy	O	B0	A0	F5	G7	K4	M0	M3	M4
V // V-K	-2	-1	0	1	2	3	4	5	6
0	0,10	1,00	3,35	6,28	11,82	22,25	39,94	70,70	125,14
1	0,06	0,63	2,11	3,96	7,46	14,04	25,20	44,61	78,96
2	0,04	0,40	1,33	2,50	4,71	8,86	15,90	28,14	49,82
3	0,02	0,25	0,84	1,58	2,97	5,59	10,03	17,76	31,43
4	0,02	0,16	0,53	0,99	1,87	3,53	6,33	11,20	19,83
5	0,01	0,10	0,33	0,63	1,18	2,23	3,99	7,07	12,51
6	0,01	0,06	0,21	0,40	0,75	1,40	2,52	4,46	7,90
7	0,00	0,04	0,13	0,25	0,47	0,89	1,59	2,81	4,98
8	0,00	0,03	0,08	0,16	0,30	0,56	1,00	1,78	3,14
9	0,00	0,02	0,05	0,10	0,19	0,35	0,63	1,12	1,98
10	0,00	0,01	0,03	0,06	0,12	0,22	0,40	0,71	1,25

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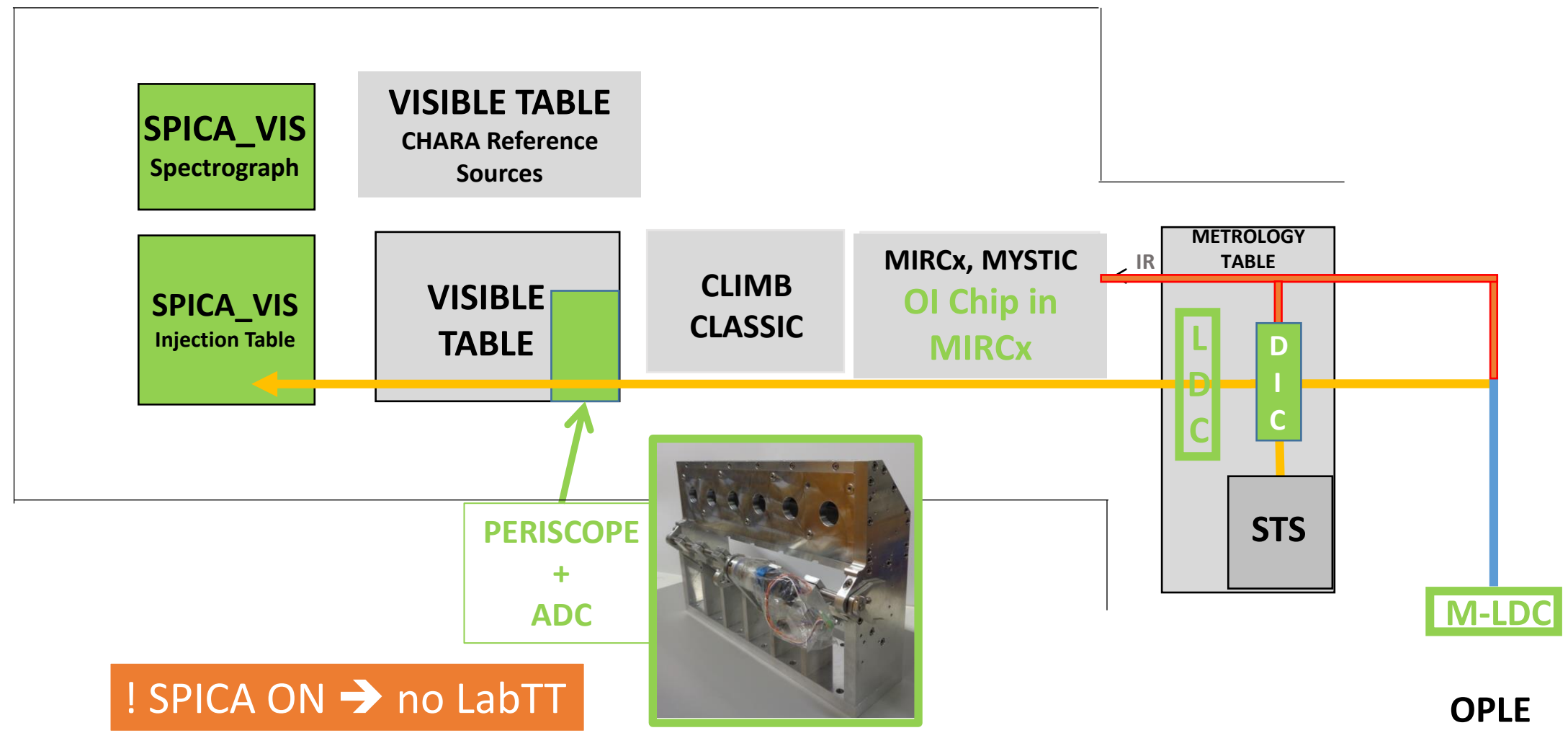
DWARFS

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SpTy	O	B0	A0	F5	G7	K4	M0	M3	M4
V // V-K	-2	-1	0	1	2	3	4	5	6
0	0,24	1,09	3,16	6,72	11,79	20,68	36,41	62,26	106,46
1	0,15	0,69	1,99	4,24	7,44	13,05	22,97	39,28	67,17
2	0,10	0,44	1,26	2,68	4,69	8,23	14,49	24,79	42,38
3	0,06	0,27	0,79	1,69	2,96	5,20	9,15	15,64	26,74
4	0,04	0,17	0,50	1,07	1,87	3,28	5,77	9,87	16,87
5	0,02	0,11	0,32	0,67	1,18	2,07	3,64	6,23	10,65
6	0,02	0,07	0,20	0,42	0,74	1,30	2,30	3,93	6,72
7	0,01	0,04	0,13	0,27	0,47	0,82	1,45	2,48	4,24
8	0,01	0,03	0,08	0,17	0,30	0,52	0,91	1,56	2,67
9	0,00	0,02	0,05	0,11	0,19	0,33	0,58	0,99	1,69
10	0,00	0,01	0,03	0,07	0,12	0,21	0,36	0,62	1,06

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8	0,01	0,03	0,08	0,17	0,30	0,52	0,91	1,56	2,67
9	0,00	0,02	0,05	0,11	0,19	0,33	0,58	0,99	1,69
10	0,002	0,011	0,032	0,067	0,118	0,207	0,364	0,623	1,065

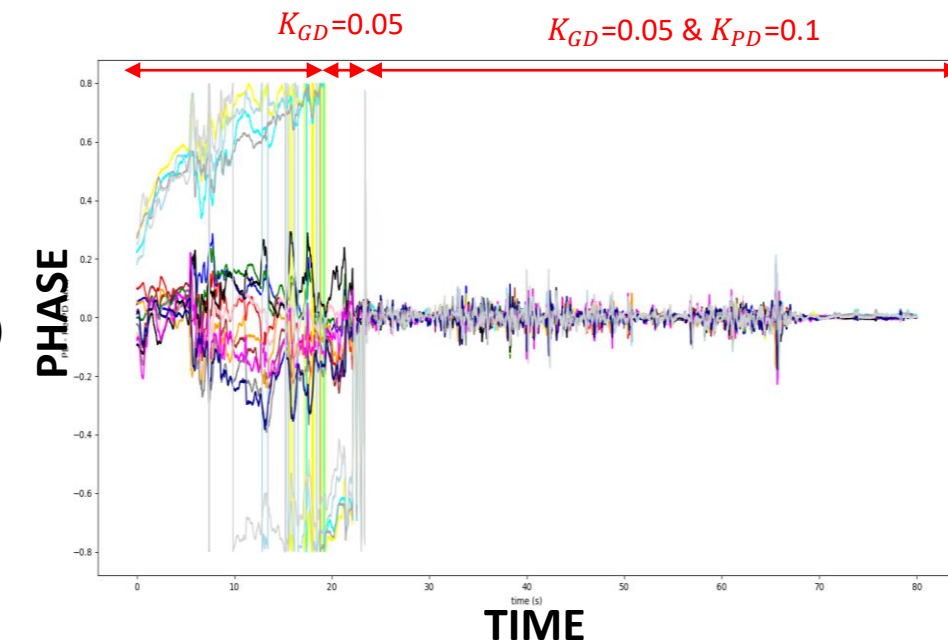
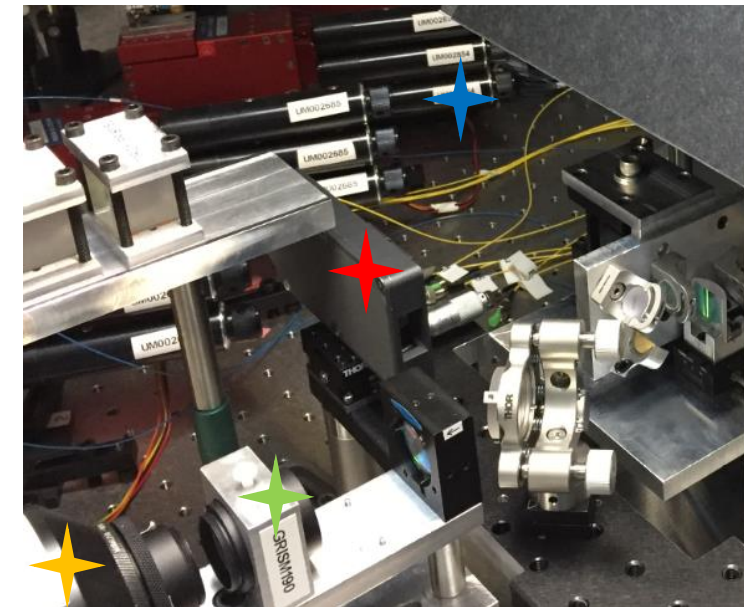
GIANTS

SPICA in the lab



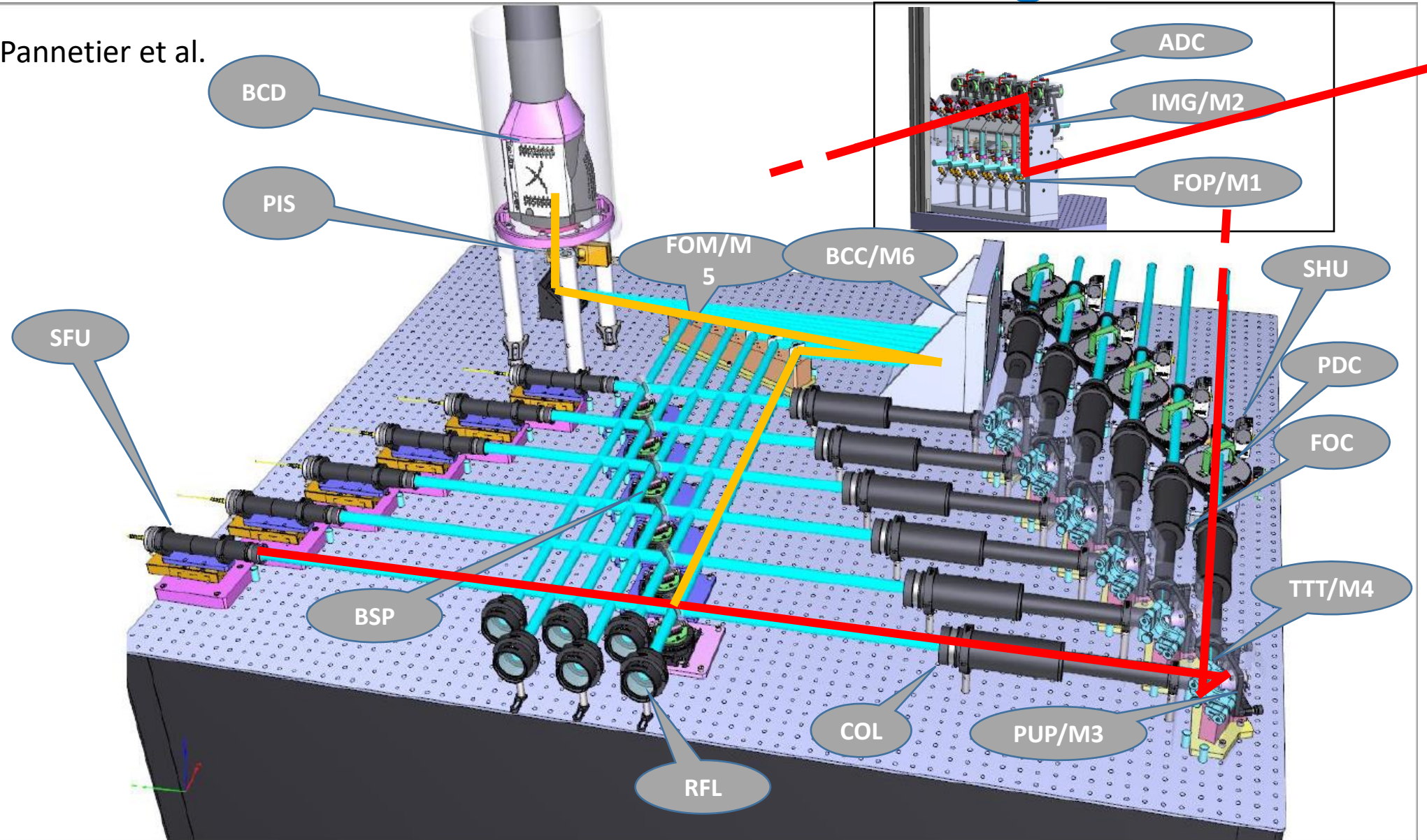
SPICA-FT

- Fibered IO device performing a **6T ABCD H band combiner**, using the **MIRCx fiber injection systems**, the **MIRCx spectrograph**, and the **CRedOne detector**.
- Specifications:
 - Exposures from 20ms (GD) to 200ms (GD+PD) in the visible, $\lambda/8$ rms
 - Goal: reaching a few seconds of integration time
- First light in January 2020 (5T, GD only)
- Optimisation of the GD+PD loops on the testbed in Nice – State Machine
- 6T fringe search operational
- Ongoing activities
 - On-sky tests, also including the new OPLE control system (remote and May?)
 - 3rd generation of the integrated optics chip: correction of internal instrumental CP, optimisation of the splitter function (summer 2021)
 - Addition of a predictive filtering (End of 2021 ?)

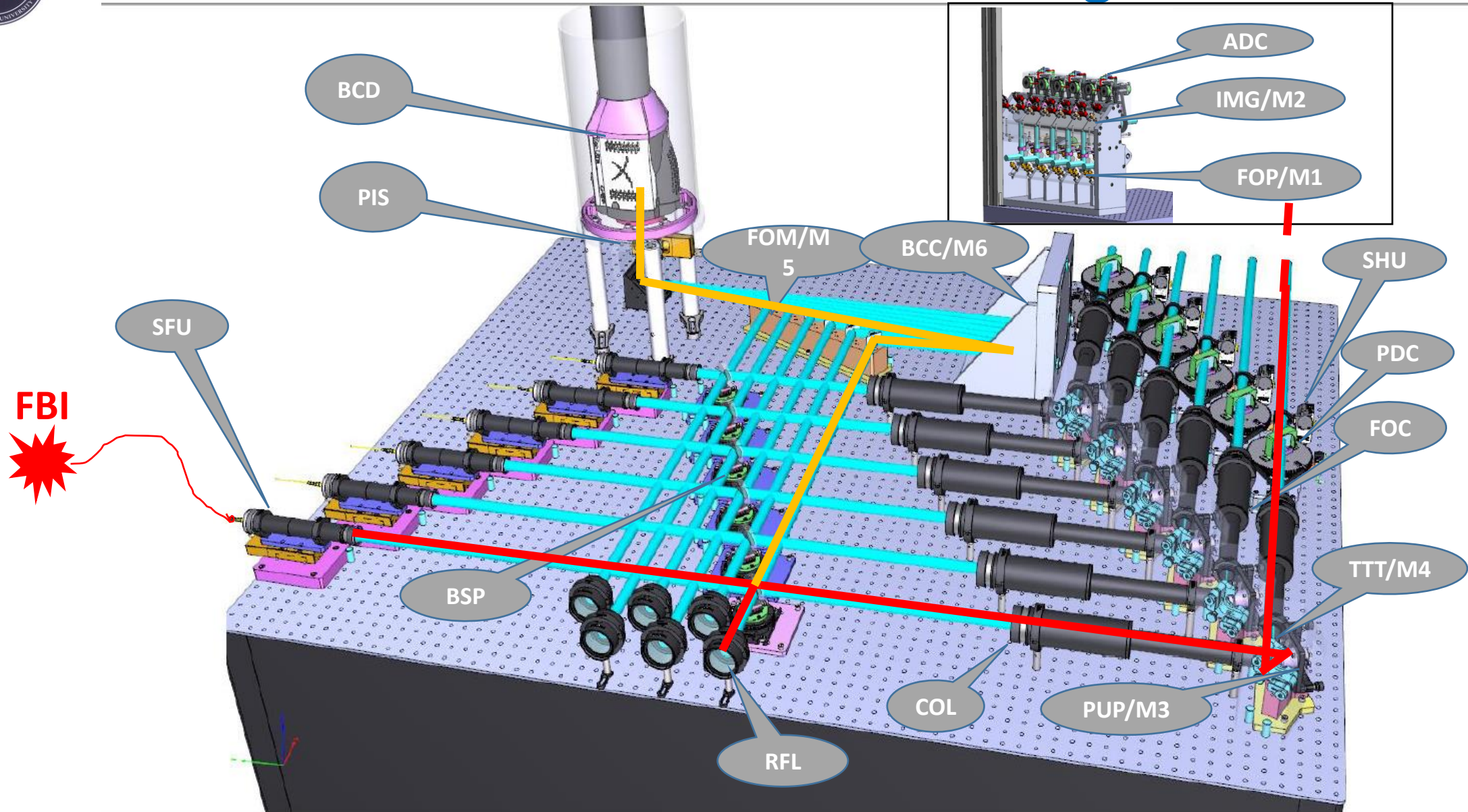


SPICA-VIS General design

SPIE 2020, Pannetier et al.



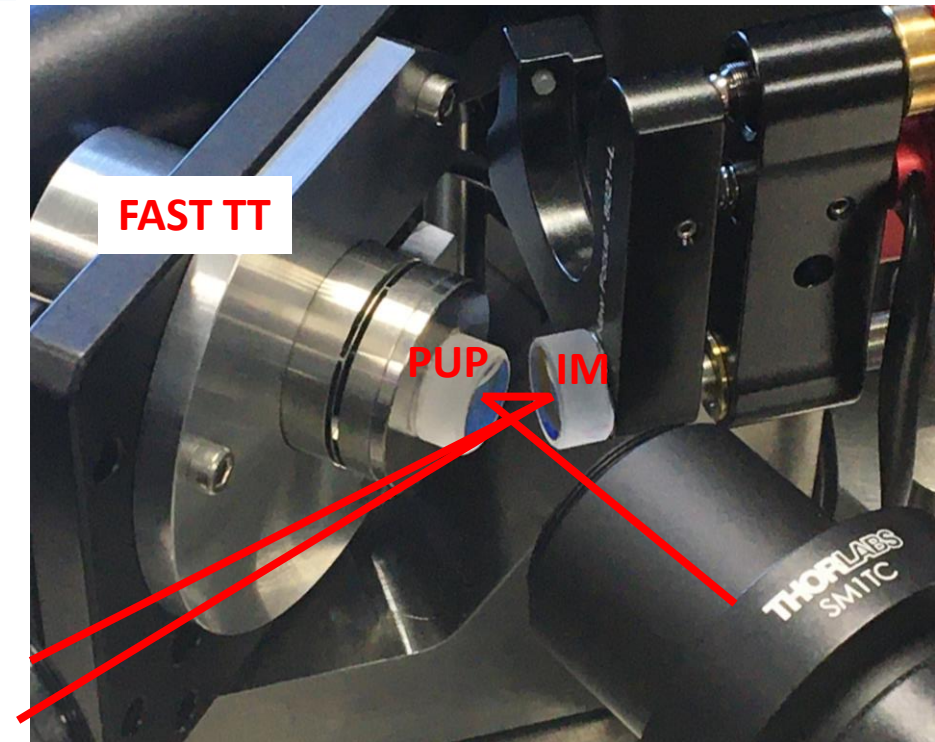
SPICA-VIS General design



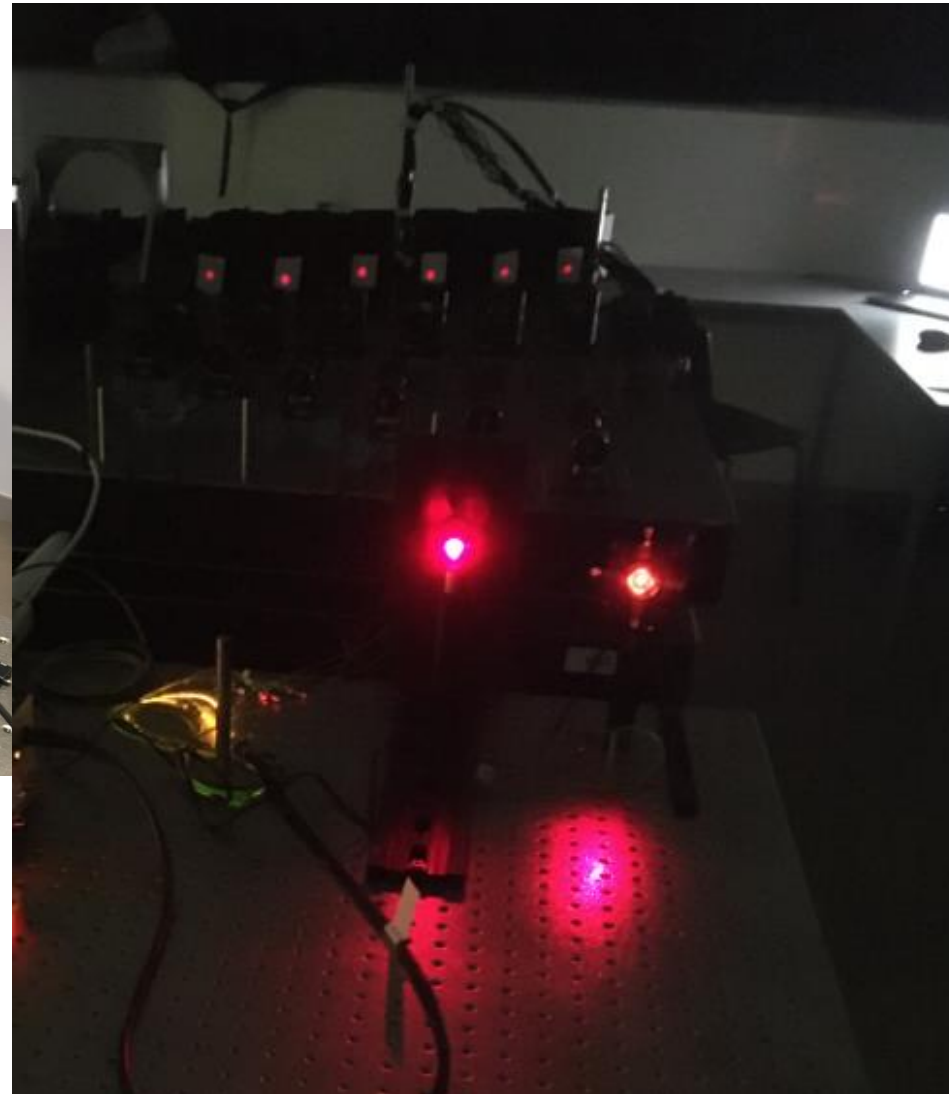
SPICA-VIS injection table

Every thing has been designed to optimize the injection in the fibres:

- Optical quality and high performance coatings
- Alignment in image and pupil plane control
- Fast tip/tilt correction to compensate the partial AO correction
- ADC for a good coupling all over the spectral band
- Fibre Back Illumination to avoid a time-expensive fibre explorer solution



SPICA-VIS injection table “*first light*” 10 Mar 2021



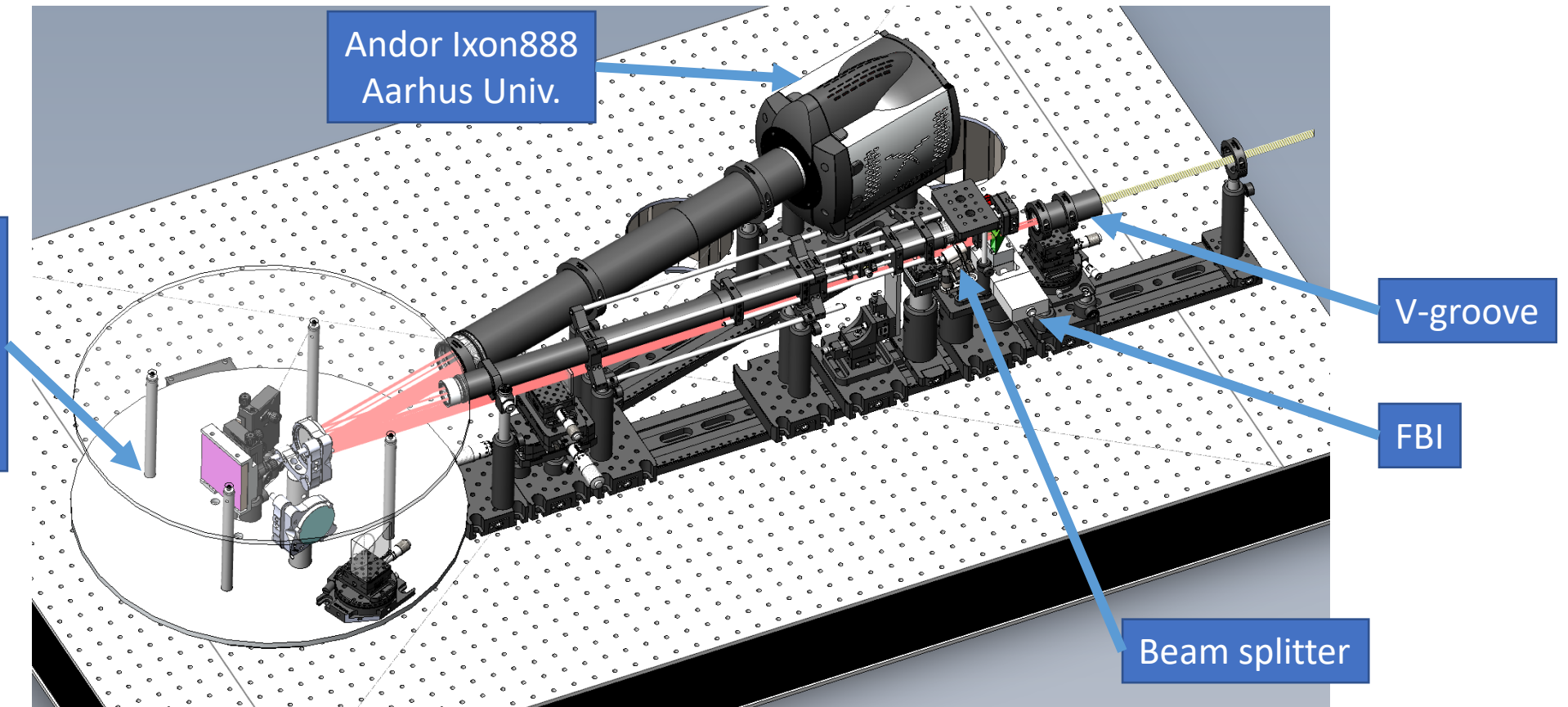
6 spots on the
FOC locations

V-groove Fiber
Back Illumination

Credit J. Dejonghe

SPICA-VIS spectrograph table

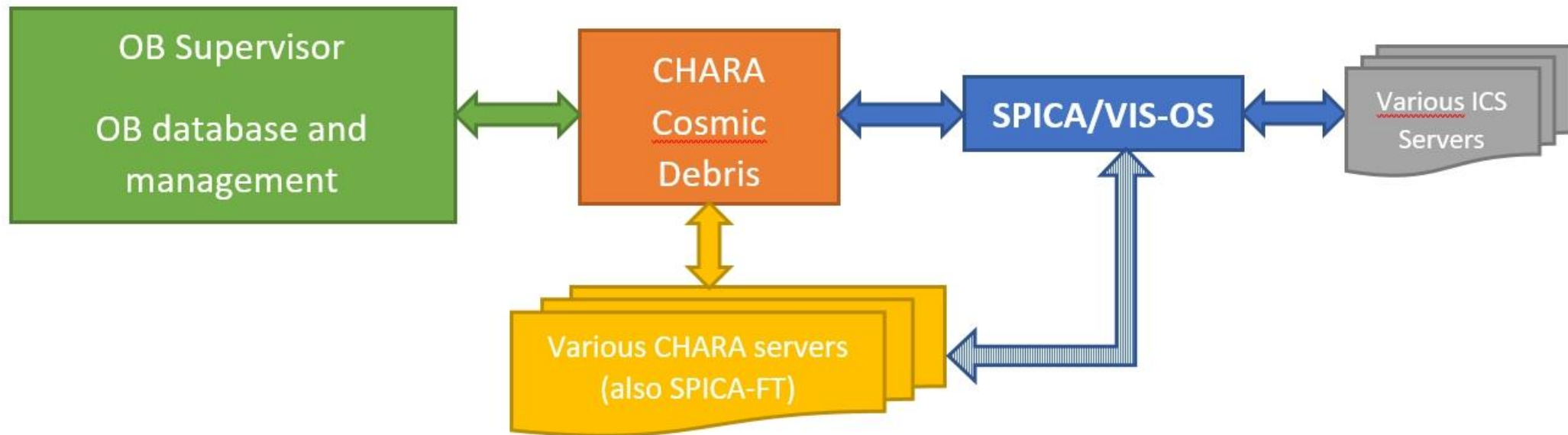
At the output of the V-groove, separation of photometric and interferometric channel
 Anamorphosis of the interferometric channel and fringes in the image plane after the dispersion
 Dispersion of the photometric channel in the 'pupil' plane
 Optical design by CP, mechanical design by JD



SPICA-VIS Observing Software

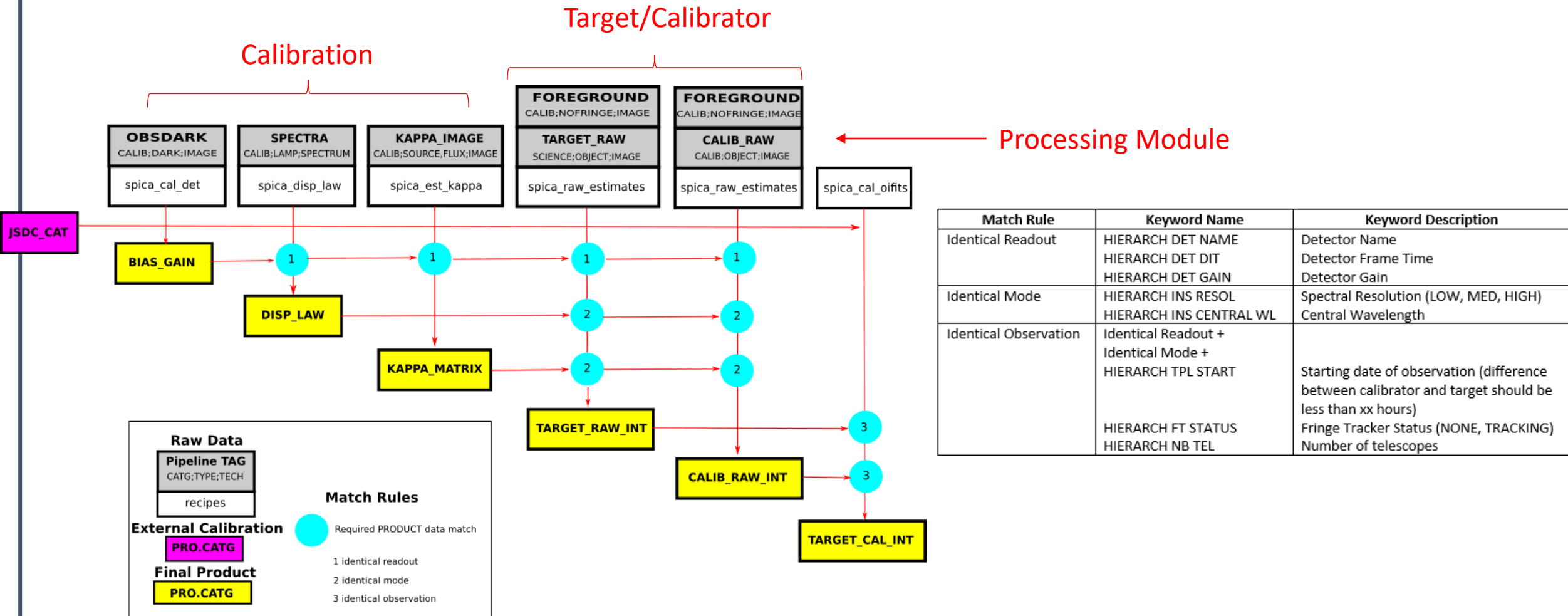
Global architecture

- CHARA Cosmic Debris is the highest client (including SPICA-FT server)
- OB supervisor for observation information (JMMC A2P2 principle adapted to CD)
- OS observation information from CD
- OS control of all SPICA-VIS devices thru servers



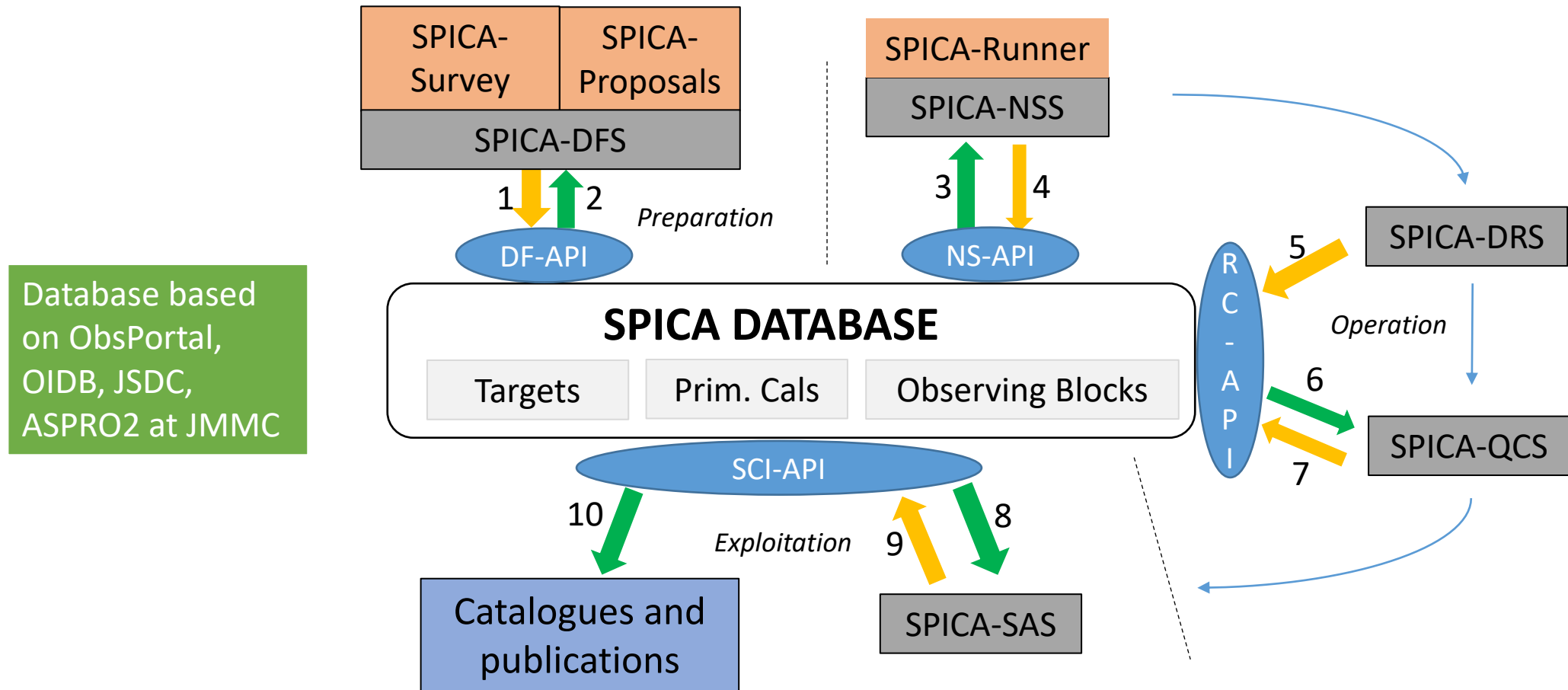


SPICA-VIS DRS (heritage from VLT/MATISSE & FRIEND)





SPICA-VIS Science operation (JMMC collaboration)

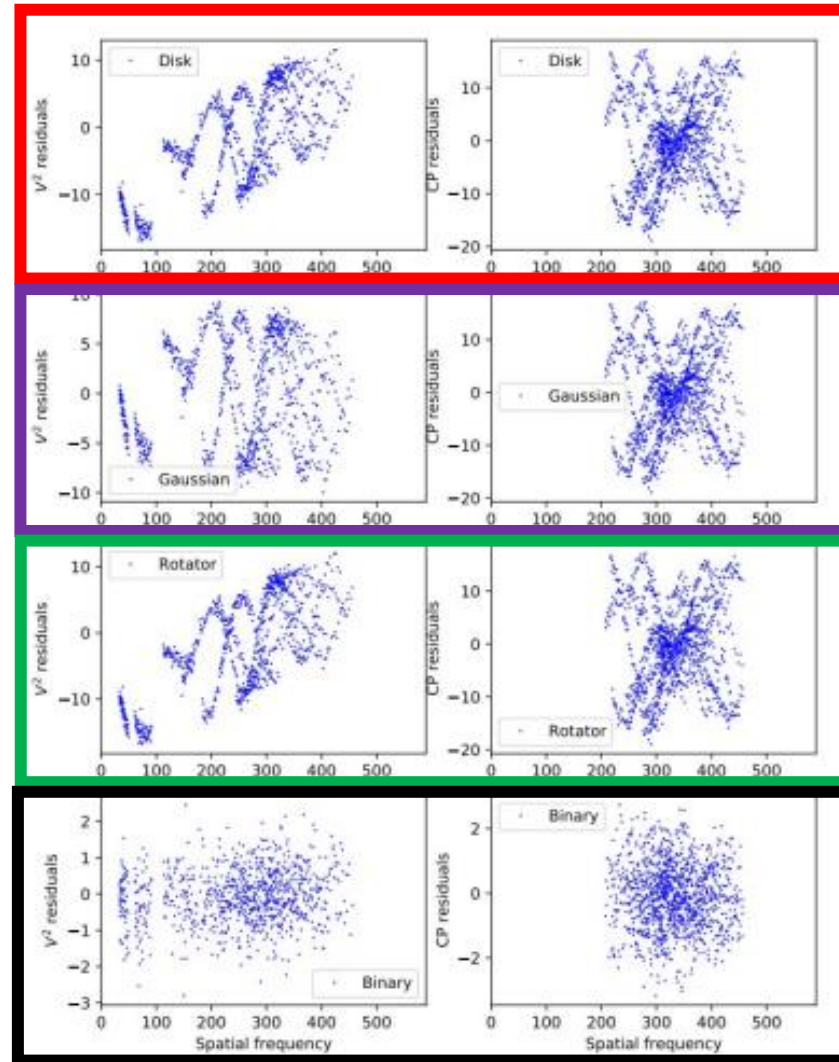


SPICA-VIS Quality Control

Automatic fit at the end of the night on all targets and calibrators. 4 models:

1. Uniform disk
2. Uniform disk + gaussian
3. Elongated disk (fast rotators)
4. Binary using CANDID (Antoine Mérand)

Computations made by D. Salabert, using ASPRO2 (SPICA) and the SPICA autofitting function.



residual V^2 residual CP

	A	B	C
1	<u>QCS UD DIAM</u>	0.373480	
2	<u>QCS UD ERRDIAM</u>	0.001192	
3	<u>QCS UD REDCHI2</u>	59.185	
4	<u>QCS UD NBD OF</u>		2099
5	<u>QCS GAUSS DIAM</u>	0.343815	
6	<u>QCS GAUSS ERDDIAM</u>	0.002100	
7	<u>QCS GAUSS FWHM</u>	16.635100	
8	<u>QCS GAUSS ERRFWHM</u>	165135603805.865784	
9	<u>QCS GAUSS FRATIO</u>	0.064068	
10	<u>QCS GAUSS ERRFRATIO</u>	0.003919	
11	<u>QCS GAUSS REDCHI2</u>	47.330	
12	<u>QCS GAUSS NBD OF</u>		2097
13	<u>QCS ROTATOR MAJORDIAM</u>	0.375595	
14	<u>QCS ROTATOR ERRMAJORDIAM</u>	0.002086	
15	<u>QCS ROTATOR ELONG</u>	1.012273	
16	<u>QCS ROTATOR ERRELONG</u>	0.010056	
17	<u>QCS ROTATOR PA</u>	75.102742	
18	<u>QCS ROTATOR ERRPA</u>	20.143640	
19	<u>QCS ROTATOR REDCHI2</u>	59.141	
20	<u>QCS ROTATOR NBD OF</u>		2097
21	<u>QCS BINARY DIAM1</u>	0.350401	
22	<u>QCS BINARY ERDDIAM1</u>	0.000160	
23	<u>QCS BINARY X</u>	1.414643	
24	<u>QCS BINARY ERX</u>	0.000448	
25	<u>QCS BINARY Y</u>	1.413492	
26	<u>QCS BINARY ERY</u>	0.000441	
27	<u>QCS BINARY FRATIO</u>	4.890021	
28	<u>QCS BINARY ERRFRATIO</u>	0.013444	
29	<u>QCS BINARY REDCHI2</u>	0.823	
30			

Database feeding



SPICA planning

