Interferometry without delay lines – the promise of Quantum Methods

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The program today

- The Quantum Memory perspective
- The benefits of high spectral resolution
- Putting it together microarcsecond arrays

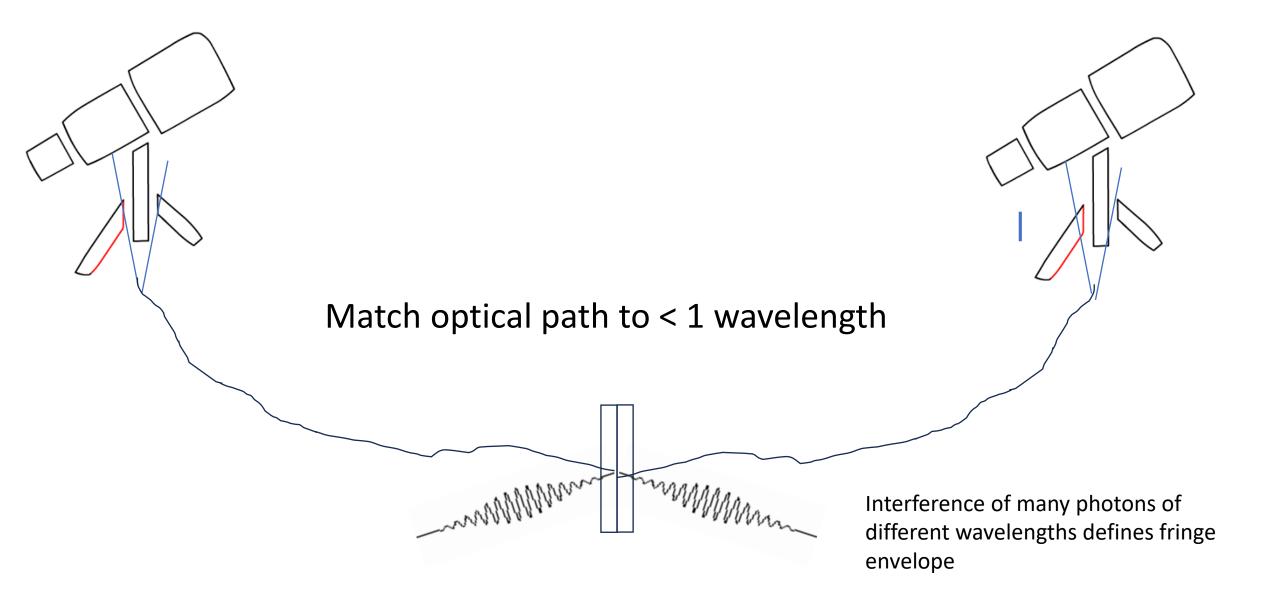
Different types of quantum memory

• Quantum Memory for Photonic Delay and Restoration ("Delay Line")

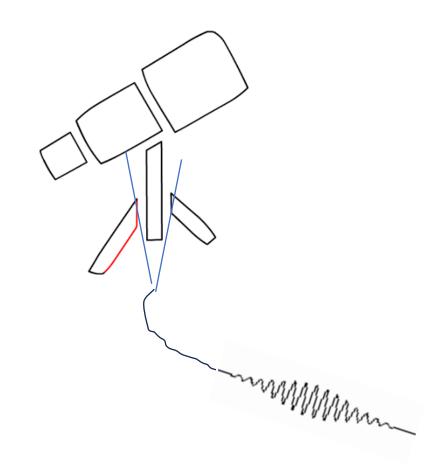
• Quantum Memory That Converts Photons into Qubits

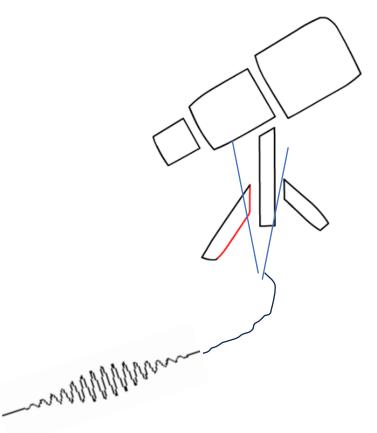
- Captures phase and coherence properties of a photon wavefront
- Photon no longer exists as propagating wave packet, but also not "measured"
- Quantum memory captures "probability" of a photon
- Qubit now "quantum computing ready" including for entanglement with another qubit

Photon capture – classical fringes view



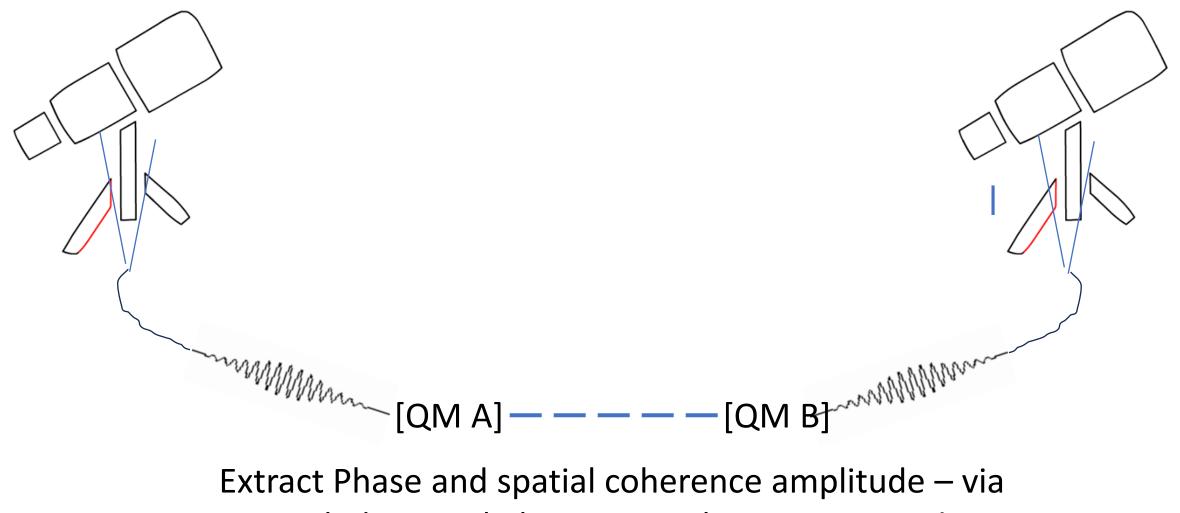
Wavefront capture – wave packet view





Interference of many "possible" wavelengths for a single photon defines coherence envelope

Wavefront capture – by quantum memory



entangled comm link.....one photon at a time!

Quantum memory – tuning fork analog

mann

The tuning fork analogy – phase between packets independent of where sampled within packet mm

Wavefront capture – wavefront -> quantum memory

mmmm QM A

Each QM acquisition has an associated temporal window

QM B

-www.www.

Wavefront capture – wavefront -> quantum memory

Match optical path to ~coherence length ~ R = $\lambda/\Delta\lambda$ wavelengths

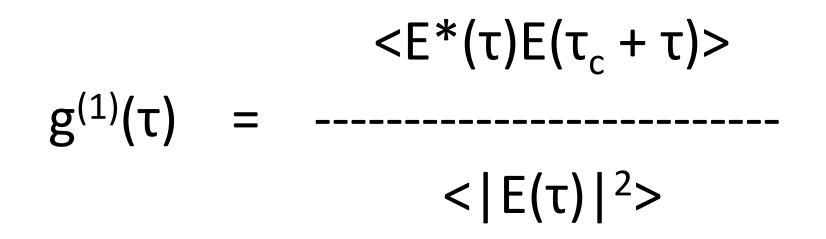
QM A

mmmm

-mannan

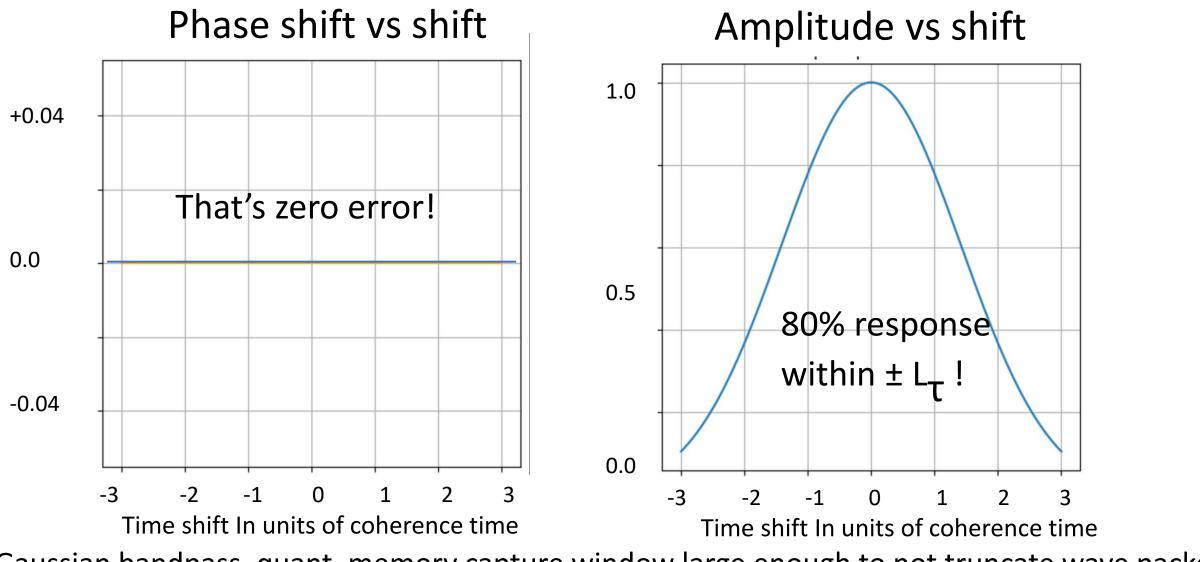
QM B

First order Glauber coherence function – describes interference visibility and phase preservation



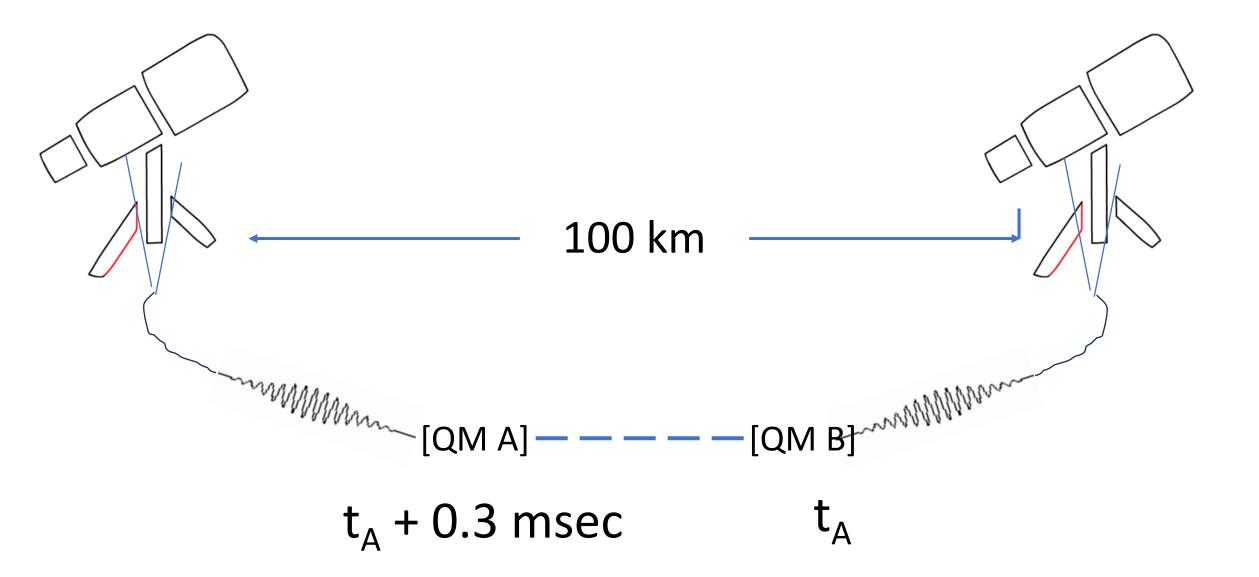
Normalized correlation function of the electric field at two space-time points.

How well can we know phase and amplitude?



*Gaussian bandpass, quant. memory capture window large enough to not truncate wave packet

Replace optical delay Δx with time delay $\Delta t = \Delta x / c$



Did I mention....

- Quantum memory and quantum communication protocols require narrow spectral bandwidth
- Some of the attractive features described here arise from that

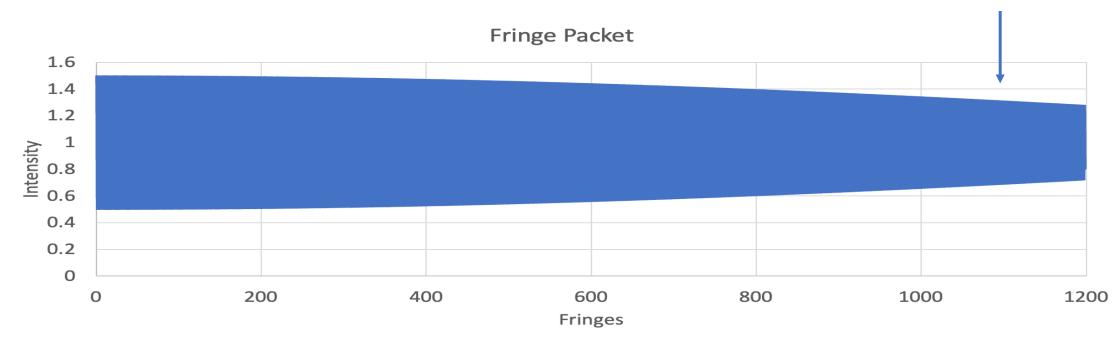
Suggest spectral parameters for discussion

- Wavelength
 - Adaptive optics performance for visible light is currently poor
 - Thermal emission > 2 μ m will not play gracefully with quantum memory
 - Select 1.6 μm as nominal spectral band
- Bandwidth
 - Quantum favors narrow bandwidth
 - Select bandwidth R = $\lambda/\Delta\lambda$ = 20000
 - May be workable R for quantum devices
 - Spectral multiplexing at this resolution already used in astronomy*

* E.g. Keck HiRes, cross-dispersed echelle with each waveband falling on a different detector pixel. We need – a 2-d array of quantum memories

Suppression of visibility 40% due to R=20000

Visualization R=20000



OPD = 6 mm

Can we "point blind" to coherence envelope?

- Wavefront FWHM (Gaussian) = 2.6 cm
- Purely GPS accuracy 1-2 mm (Earth-based inertial coord system)
- What Earth modeling adds: precession, nutation, sidereal to rotation
 - Offers ~ 1mm differential relative to stars
- VLBI
 - Offers less than 1mm
- Yes, can predict fringe packet zero path to a fraction of its width

Can we match time bins at different telescopes?

- A FWHM 2.6 cm fringe packet has a temporal duration of 48 ps
- A rubidium clock synchronized to GPS GPS-disciplined oscillator (GPSDO)
 - Accurate to 0.1 ps
- Yes, we can select corresponding time bins from different telescopes with confidence

Determine phase and visibility at 2+ wavelengths

Observe 2 wavelengths differing by 1/20000

- Phase variation with wavelength => OPD
- OPD allows correction of visibility to ZPD value
- OPD measure supports OPD control

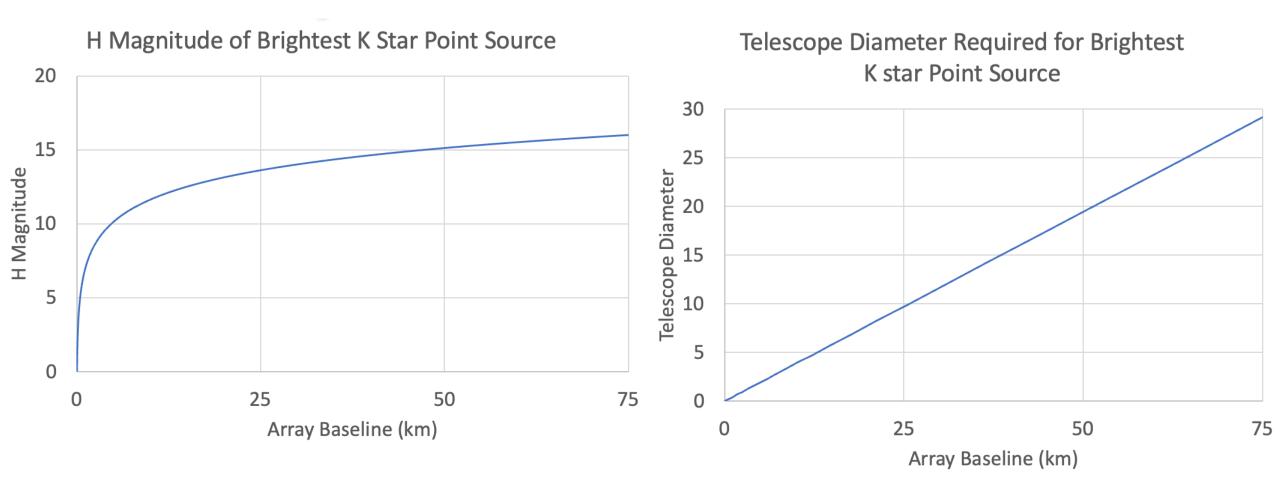
OPD dither by quantum ops over fractional wavelength Complete interferometric measurement with OPD up to ~ cm from ZPD

Conclusions on Feasibility of Eliminating Delay Lines

- Purely by metrology and timing, we can "point" an array and sample wavefronts within a fraction of the wavefront coherence length
- This can support quantitative interferometry without optical path servo

- Without servo, we would prefer to have enough signal to calibrate the spatial coherence amplitude. This requires a minimum signal (over two or more spectral bands)
- In a lower signal regime, it may still be possible to acquire fringe closure measurements one photon at a time.

Designing a Microarcsecond Array



With current interferometer efficiencies, need bootstrapping beyond ~10-20 km baseline

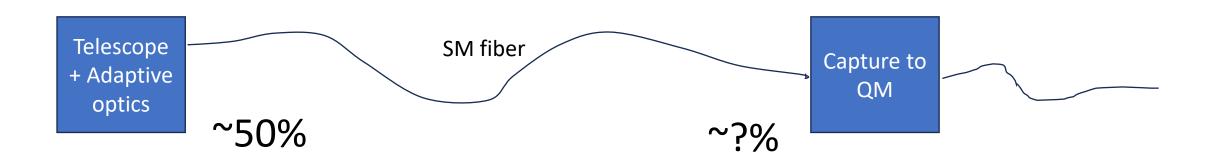
Assumptions in previous slide

- Fringe reference star is a K star observed in the H band
- Fringe tracking requirement defined by VLTI performance
 - AT Diameter = 1.8m
 - Limit fringe tracking magnitude = 9.5
 - Photon rate = 7.07E5 per second

Efficiency improvements?



Classical – 1% throughput



Next steps toward quantum interferometery

- Parameters to develop next:
 - Quantum memory characteristics
 - Quantum communications/computing support
- Efficiency budget for photon losses
- Calibration budget for derived visibilities and phases
- Astronomers observing calculator
 - Predicted S/N for standard measurements of simple targets
 - Point sources
 - Uniform disks
 - Equal amplitude binaries
 - Faint companions