



Observing Strategies and Planning Software



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Outline

- Selecting a Beam Combiner
- Selecting Telescopes and Baselines
- Selecting Calibrator Stars
- Selecting Delay Settings (POP Configuration)
- Time Needed for Observations



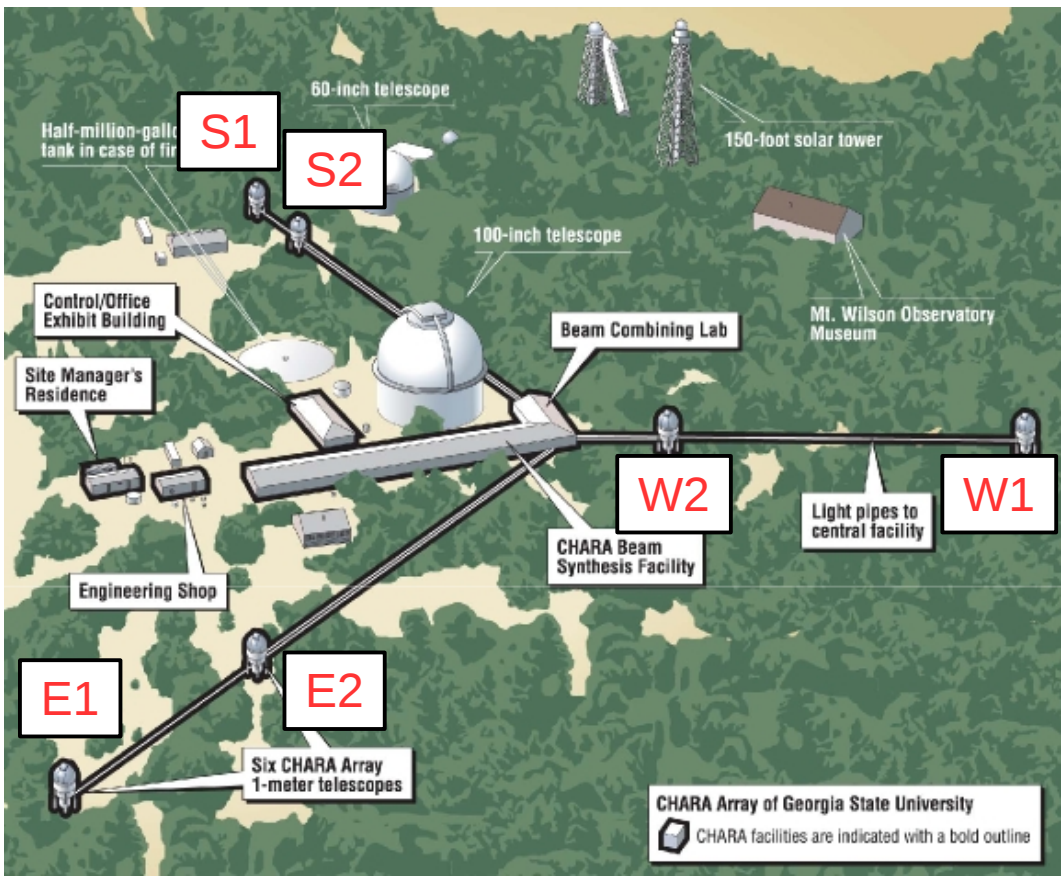
Beam Combiners

Combiner	Num Tel.	Band	Typical Mag	Best Mag	Spec. Res.	Advantages
CLASSIC	2T	H or K	7.0	8.5	Broad	Sensitivity
CLIMB	3T	H or K	6.0	7.0	Broad	Sensitivity
JouFlu	2T	K	4.5	5.0	Broad	Precision
MIRC	6T	H	4.5	6.0	40	Imaging
PAVO	2T	630-900 nm	7.0	8.0	30	Sensitivity
VEGA – HiRes	2-3T	2 bands (7nm) in 480-850 nm	4.0	5.0	30000	Spectral Res
VEGA – MedR	2-3T	2 bands (35 nm) in 480-850 nm	6.5	7.5	6000	Spectral Res

Limit for acquisition and tip/tilt tracking: $V = 10-12$ mag



Telescopes and Baselines



Baseline	Length (m)
E1-S1	331
W1-E1	314
E1-S2	302
E2-S1	279
W1-S1	279
W1-E2	251
W1-S2	249
E2-S2	248
W2-S1	211
W2-E1	222
W2-S2	177
W2-E2	156
W2-W1	108
E2-E1	66
S2-S1	34

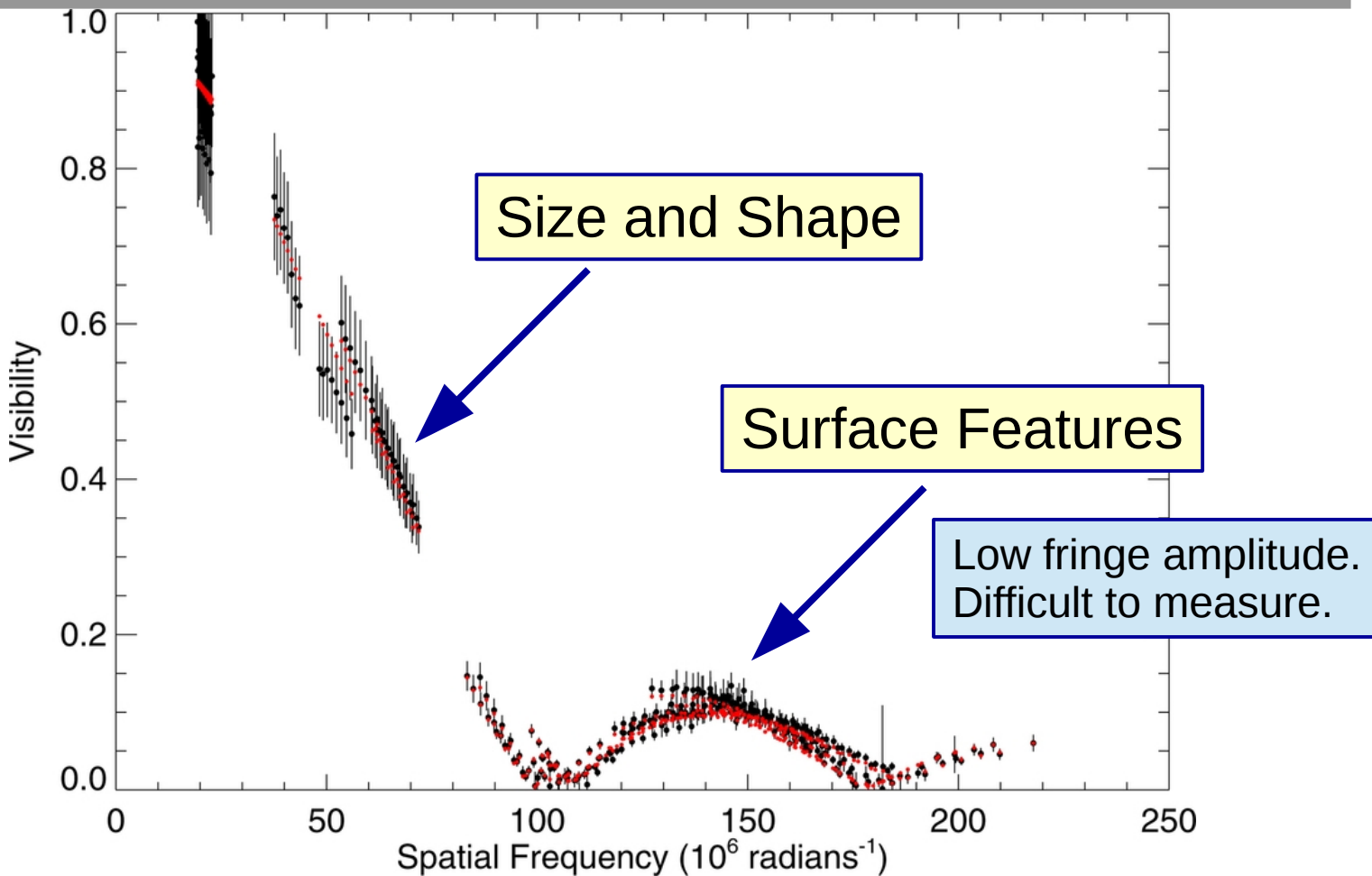


Selecting Beam Combiner and Baselines

- Angular Resolution: $0.5 \lambda / B$
 - 0.66 mas in K-band (2.13 μm)
 - 0.52 mas in H-band (1.67 μm)
 - 0.20 mas in visible at 650 nm
- Simple diameter:
 - Single baseline (two telescopes)
- Imaging complex sources: Rapid rotators, binaries, stellar surfaces
 - Multiple baseline projections
 - Sample beyond the first null (at $1.22 \lambda / B$)



Selecting Beam Combiner and Baselines





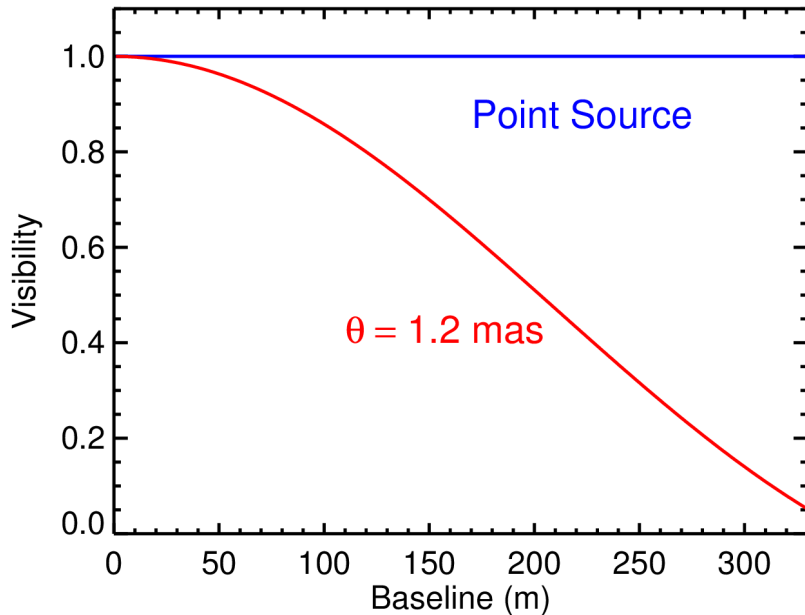
Wavelength Coverage

- Broad-band
 - Higher sensitivity for faint objects
 - Bandwidth smearing
- Spectrally dispersed visibilities
 - Increase u,v sampling by measuring fringes in different wavelength channels
 - Emission/absorption line studies
- Coherence length (width of fringe packet)
 - Sets the interferometric field of view

$$\frac{\lambda^2}{\Delta\lambda}$$



Calibrator Stars



- Unresolved point source:
 - Visibility = 1.0
- However, instrumental and atmospheric effects will cause a loss in coherence, causing a drop in the measured visibility.
- Observe unresolved calibrator stars to define the true visibility of the target.



Selecting Calibrators

- Unresolved stars or stars with a known angular diameter.
- Within 5-10 degrees on the sky from the science target.
- Within 1-2 mag in brightness from science target and similar in color.
- Avoid binary stars, rapid rotators, emission line stars.
- Minimum of two calibrators per object, three is better.
 - Discovery of unknown binaries



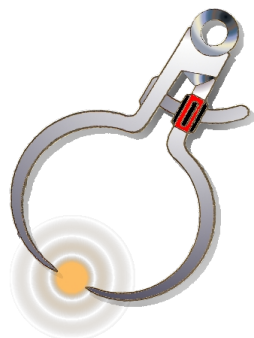
Selecting Calibrators

- **SearchCal** – developed by JMMC
 - http://www.jmmc.fr/searchcal_page.htm
- **getCal** – developed by NexSci
 - <http://nexsciweb.ipac.caltech.edu/gcWeb/gcWeb.jsp>



SearchCal

JMMC



SearchCal [c1]

File Edit Query Calibrators Interop Help

1) Instrumental Configuration

Magnitude Band :

Wavelength (H) [μm] :

Max. Baseline [m] :

2) Science Object

Name :

RA 2000 [hh:mm:ss] :

DEC 2000 [+/-dd:mm:ss] :

Magnitude (H) :

3) SearchCal Parameters

Min. Magnitude (H) :

Max. Magnitude (H) :

Scenario : Bright Faint

RA Range [mn] :

DEC Range [deg] :

Progress :

Found Calibrators (2041 sources, 1826 filtered)

Index	dist /	HD	RAJ2000	DEJ2000	vis2	LDD	UD_V	UD_H	UD_K	SpType	V	H	K
1	5.21E-6	69897	08 20 03.8602	+27 13 03.7380	0.374	0.701	0.662	0.686	0.689	F6V	5.13	3.942	3.868
2	2.975	67542	08 09 35.1816	+29 05 35.0772	0.622	0.503	0.468	0.488	0.49	G0III	6.47	4.699	4.621
3	3.383	67544	08 09 24.8645	+24 49 34.0716	0.619	0.509	0.468	0.49	0.493	G8III	7.29	4.966	4.908
4	3.595	71730	08 29 40.0634	+24 20 40.9452	0.581	0.542	0.496	0.52	0.523	K0III	7.05	4.872	4.81
5	3.977	73080	08 37 22.1112	+28 17 39.8328	0.554	0.555	0.52	0.541	0.543	G5	6.63	4.702	4.591
6	4.945		08 10 54.7320	+22 43 43.1904	0.514	0.588	0.548	0.573	0.575	K0	8.026	4.909	4.698
7	6.121	65471	07 59 42.6055	+23 10 58.4652	0.506	0.594	0.554	0.579	0.581	K0	6.92	4.73	4.562
8	6.902	75216	08 49 45.3118	+29 26 55.9824	0.534	0.581	0.529	0.557	0.561	K2III	7.38	4.868	4.712
9	7.14	63138	07 48 28.8108	+28 45 51.2748	0.519	0.592	0.542	0.568	0.572	K0III	6.86	4.694	4.605
10	7.303	75646	08 52 00.4543	+25 43 07.1004	0.568	0.553	0.504	0.53	0.534	K2III	7.54	4.983	4.834
11	7.524		08 52 09.6634	+29 51 13.3848	0.515	0.588	0.548	0.572	0.575	K0	7.08	4.742	4.631
12	7.677		08 32 54.2333	+34 23 03.2748	0.544	0.565	0.524	0.549	0.552	K2	7.52	4.934	4.798
13	7.731	75783	08 53 00.0972	+29 57 41.5296	0.564	0.549	0.509	0.533	0.536	K2	7.35	4.982	4.813
14	7.809	74198	08 43 17.1466	+21 28 06.6000	0.774	0.366	0.351	0.362	0.362	A1V	4.66	4.788	4.638
15	7.842	64092	07 53 01.0094	+22 20 04.3116	0.557	0.553	0.515	0.538	0.541	K0	7.05	4.85	4.755
16	7.845		08 03 34.1340	+20 20 18.6972	0.599	0.519	0.486	0.506	0.508	G5	7.03	4.836	4.742
17	8.769	67482	08 09 39.7601	+35 42 08.5032	0.58	0.535	0.498	0.52	0.523	K0	7.3	4.952	4.839
18	8.815	64602	07 56 01.9399	+34 22 10.4160	0.572	0.541	0.505	0.527	0.53	K0	7.57	4.972	4.827
19	8.917		07 48 06.8957	+32 51 25.0308	0.557	0.552	0.518	0.539	0.541	G5	7.204	4.952	4.548
20	10.159	60204	07 34 31.5922	+28 41 11.6808	0.519	0.583	0.547	0.569	0.571	G5	6.66	4.605	4.498
21	10.402	77694	09 04 51.4817	+24 36 18.5040	0.561	0.559	0.509	0.536	0.539	K2III	7.8	4.982	4.838

Filters

Reject stars farther than : Maximum RA Separation (mn) : Maximum DEC Separation (degree) :

Reject stars with magnitude : below : and above :

Reject Spectral Types (and unknowns) : O B A F G K M

Reject Luminosity Classes (and unknowns) : I II III IV V VI

Reject Visibility below : vis2 :

Reject Visibility Accuracy above (or unknown) : vis2Err/vis2 (%) :

Reject Variability

Reject Multiplicity

Reject Invalid Object Types

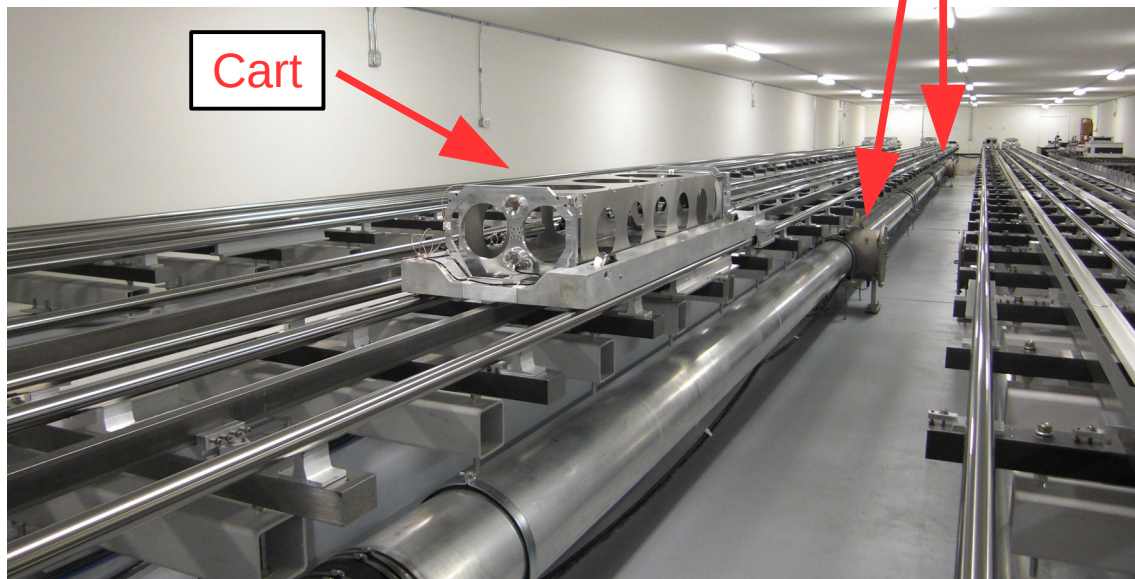
Diameter quality : Maximum chi square : Maximum relative error (%) :

searching calibrators... done.

29 M Provided by JMMC

Delay settings (sky coverage)

- Delay settings to equalize optical path length
 - Fixed increments: PoPs
 - Variable delay: carts



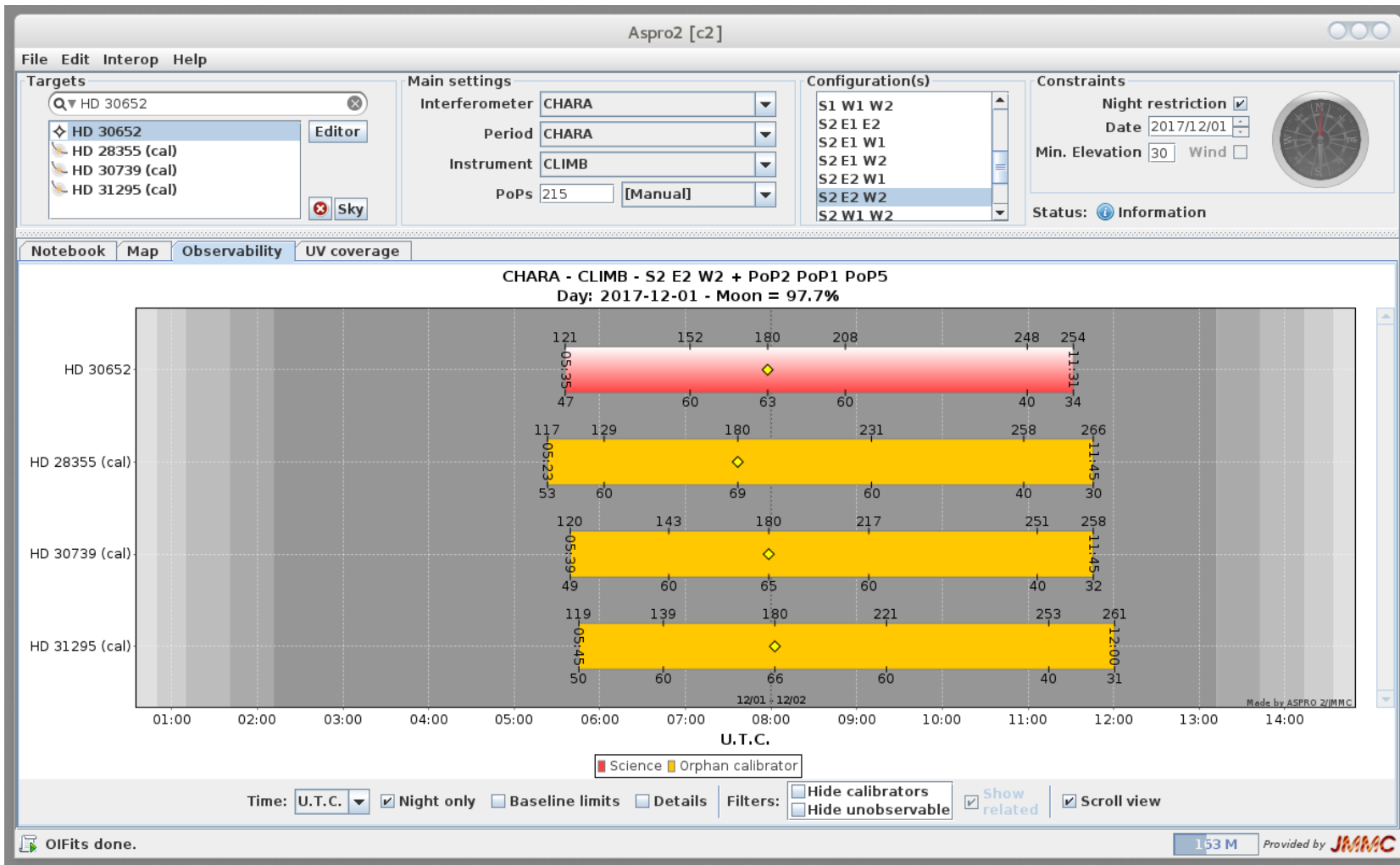


Planning Software

- **ASPRO2** – developed by JMMC
 - http://www.jmmc.fr/aspro_page.htm
- **CHARA_PLAN2** – developed by CHARA
 - http://www.astro.gsu.edu/~theo/chara_reduction/index.html

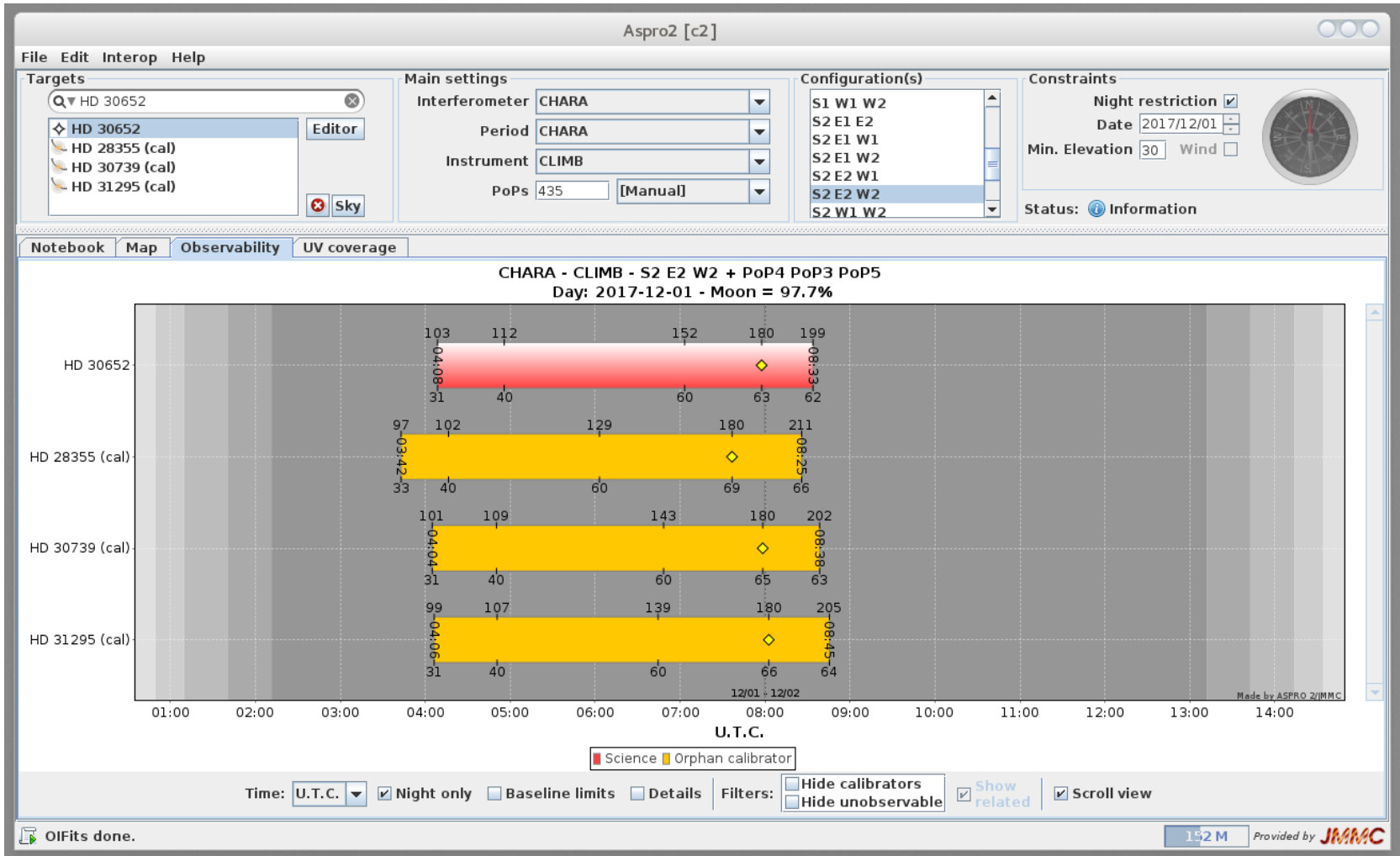


ASPRO 2





ASPRO 2





ASPRO 2



Aspro2 [c1]

File Edit Interop Help

Targets
Q HD 56537
HD 30652 Editor
HD 28355 (cal)
HD 30739 (cal)
HD 31295 (cal)
Sky

Main settings
Interferometer: CHARA
Period: CHARA
Instrument: CLIMB
PoPs: 215 215

Configuration(s)
S1 W1 W2
S2 E1 E2
S2 E1 W1
S2 E1 W2
S2 E2 W1
S2 E2 W2
S2 W1 W2

Constraints
Night restriction
Date: 2017/12/01
Min. Elevation: 30 Wind
Status: Information

Notebook Map Observability UV coverage OIFits viewer

Instrument mode
H
Atmosphere quality: Average
U-V range to plot (m): 347.20
Sampling Periodicity (min): 40
Total Integration time (s): 300
HA min: -2.37
HA max: 3.58
 Plot rise/set uv tracks
 Underplot a model image
Plot what ... AMP
 Compute OIFits data
 Add error noise to data
 Use inst. & cal. error bias

CHARA - CLIMB - S2 E2 W2 + PoP2 PoP1 PoP5
Day: 2017-12-01 - Source: HD 30652

U (m) - North
V (m) - East

U (mÅ)
V (mÅ)

Made by ASPRO 2/JMMC

117 M Provided by JMMC

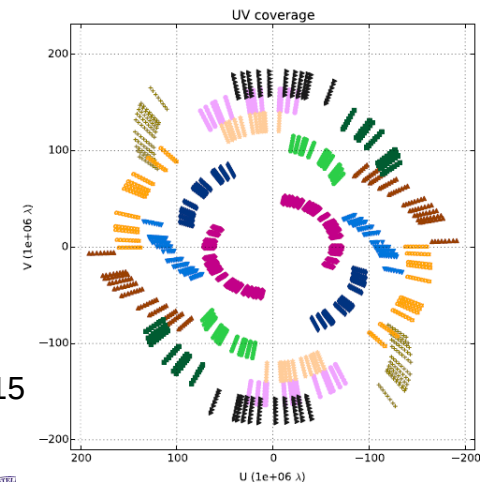


How much time is needed?

- Calibration Strategies:
 - Cal1 – Obj – Cal2 – Obj – Cal1 ...
 - Cal1 – Obj – Cal2 – Cal1 – Obj – Cal2 ...
- Time to collect data on single object (star acq. + data)
 - Seeing and brightness dependent
 - Fast instruments (CLASSIC, CLIMB, PAVO, JouFlu):
 - 5 – 10 minutes
 - VEGA: 5 – 20 minutes
 - MIRC: 45 – 60 minutes
- Cal-Sci-Cal will take between 30 – 120 min

How much data is needed?

- Diameters – Several brackets of data per baseline on two separate nights.
- Binaries – Minimum of three brackets or observations on at least three baselines.
- Imaging – Many brackets on multiple baselines during the night to fill in the sky coverage.



Kloppenborg et al. 2015



Do observations already exist?

- OI Database
- Query and download data (OIFITS)
- CHARA observation logs for Classic, CLIMB, VEGA only

JMMC - OI DB - Home Search Submit new data Help - Sign in

Filters

Position: GJ 581 Radius: 2 arcmin Date of observation: after YYYY-MM-DD before YYYY-MM-DD

Instrument: Any Instrument Wavelength range: any value

Collection: Any Collection DataPI name: Any DataPI

Data reduction level: L0, L1, L2, L3 Availability: Public Restricted All

25 rows max. per page, sorted by Instrument descending. Search Reset

Results

Meta-data will try to follow VO4OI proposal and Ivoa:ObsCore document (get metadata description in the associated doc)
33 observations from 1 oifits files (0 private)

Page 1 / 2 Next Last

Results for

```
SELECT ALL * FROM oidb AS t WHERE ( CONTAINS(POINT('ICRS', t.s_ra, t.s_dec), CIRCLE('ICRS', 229.8617625, -7.7222806, 0.033333333333333333))=1 ) ORDER BY instrument_
```

(Edit query)

target_name	access_url	t_min	instrument_name	wlen_min	wlen_max	nb_channels	datapi
HIP_74995	-	2008-05-16T09:38:52	CLASSIC	1.96000000	2.31000000	-	Baines
HIP_74995	-	2010-03-30T08:09:35	CLASSIC	1.53000000	1.82000000	-	Boyajian
HIP_74995	-	2010-03-30T08:31:12	CLASSIC	1.53000000	1.82000000	-	Boyajian
HIP_74995	-	2010-03-30T09:44:38	CLASSIC	1.53000000	1.82000000	-	Boyajian
HIP_74995	-	2010-03-30T10:13:26	CLASSIC	1.53000000	1.82000000	-	Boyajian

<http://oidb.jmmc.fr/index.html>



LESIA



Observatoire de la COTE d'AZUR

UNIVERSITY OF EXETER



On the night of observation

- Observations will be carried out by CHARA staff
- Visitors are encouraged to travel to the Array to participate in the observations
 - Real-time input from PI on decisions that could impact the science objectives and priorities



Example 1: Angular Diameter

- **Exoplanet host star: 55 Cnc** (von Braun et al. 2012)
 - $V = 5.6$ mag, $K = 4.0$ mag
 - $\theta = 0.71$ mas
 - CLASSIC H/K-band, 2T combiner
 - Baselines longer than 250 m (ang res ~ 0.69 mas at H)
- **Asteroseismic target: HD 182736** (Huber et al. 2012)
 - Subgiant showing solar-like oscillations
 - $V = 7.0$ mag, $K = 5.0$ mag
 - $\theta = 0.44$ mas
 - PAVO R-band, 2T combiner
 - Baselines longer than 150 m (ang res ~ 0.43 mas at R)

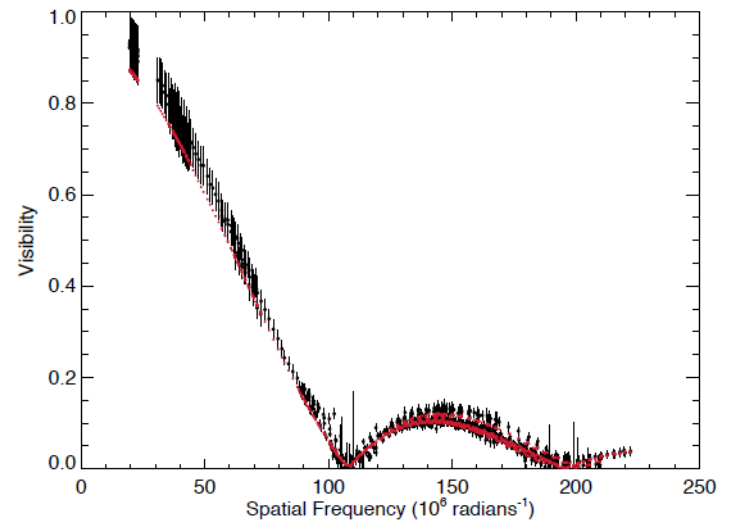


Example 2: Binary Star

- **Wolf Rayet Binary - WR 137** (Richardson et al. 2016)
 - $V = 7.9$ mag, $H = 6.8$ mag, $K = 6.2$ mag
 - sep $\sim 4 - 10$ mas, flux ratio $\sim 60\%$
 - **CLIMB H/K-band, 3T combiner** [2D coverage, faint target]
 - **Inner triangle (S2-E2-W2), baselines 140-250 m (~ 0.7 mas)**
- **Faint Cepheid companion – V1334 Cyg** (Gallenne et al. 2015)
 - $V = 5.9$ mag, $H = 4.7$ mag, $K = 4.5$ mag
 - sep ~ 8 mas, flux ratio $\sim 3.1\%$
 - **MIRC H-band, 6T combiner**
 - **All 15 baselines 34-331 m** [high precision closure phases!]
 - **8 spectral channels** [longer coherence length]

Example 3: Imaging

- Spotted giant star - sig Gem (Roettenbacher et al. 2017)
 - $V = 4.3$ mag, $H = 1.8$ mag
 - $\theta = 2.4$ mas
 - MIRC H-band, 6T combiner [Good uv coverage on sky]
 - All 15 baselines (34-331 m)
 - Sample 2nd lobe in visibility curve!





Guide to planning observations available on the CHARA website:

<http://www.chara.gsu.edu/observers/planning-an-observation>